

# Phenomenology of Extra Dimensions

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- Large
- TeV
- ★ • Warped - Davoudiasl, JLH, Rizzo  
PRL '00, PLB '00, hep-ph/0006041  
10006097

## Consequences for:

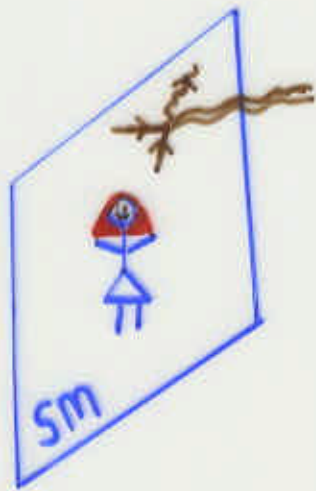
- Cosmology
- Astrophysics
- ★ Collider physics
- ★ Precision measurements
- Flavor physics
- Gravity tests

Hewett  
ICHEP2000

# Large Extra Dimensions

Arkani-Hamed,  
Dimopoulos, Dvali

Motivation: Solve the hierarchy problem by removing it!



Gauss' Law:

$$M_{Pl}^2 = V_n M_*^{2+n} ; \quad V_n \sim R^n$$

$M_*$  = Fundamental Planck scale  
in the bulk

$$\approx \text{TeV}$$

$n=1$	$R=10^{11}$ m	<b>Excluded!</b>
2	0.4 mm	$m_c = 1/R = 5 \times 10^{-4}$ eV
4	$10^{-5}$ mm	20 keV
6	30 fm	7 MeV

$n=2$ : Low  $M_*$  disfavored by Astro/Cosmology

(i) Supernova Cooling  $\rightarrow M_* \gtrsim 30$  TeV

$NN \rightarrow NN + G_n$  can cool supernova too rapidly

Cullen + Perelstein  
Barger et al  
Hanhart et al

(ii)  $\gamma$ -ray flux  $\rightarrow M_* \gtrsim 110$  TeV

$\nu\bar{\nu} \rightarrow G_n \rightarrow \gamma\gamma$  produces too many soft  $\gamma$ 's

Hall + Smith

## Bulk Metric: Linearized Quantum Gravity

$$G_{AB} = \eta_{AB} + \frac{h_{AB}(x^M, x^a)}{m_*^{n/2+1}}$$

$$\begin{aligned} A &= 0, \dots, 4+n \\ \mu &= 0, 1, 2, 3 \\ a &= 1, \dots, n \end{aligned}$$

with interactions

$$S_{\text{int}} = \frac{-1}{m_*^{n/2+1}} \int d^4x d^n x^a h_{AB}(x^M, x^a) T_{AB}(x^M, x^a)$$

Induced wall metric:  $G_{\mu\nu}(x^M, x^a=0)$

SM on wall:  $T_{AB} = \eta^{\mu}_A \eta^{\nu}_B T_{\mu\nu} \delta(x^a)$

Interactions on wall: Decompose  $h_{AB}$

impose unitary gauge

integrate  $S_{\text{int}}$  over  $d^n x^a$  via  $\delta(x^a)$

Bulk fields expand into Kaluza-Klein towers upon compactification

$$\Phi(x^M, x^a) = \sum_{n=0}^{\infty} \phi^{(n)}(x^M) e^{inx^a/R} \cdot \frac{1}{\sqrt{V_n}}$$

$\delta^2 \Phi = 0$  gives massless mode in  $4+n$  D

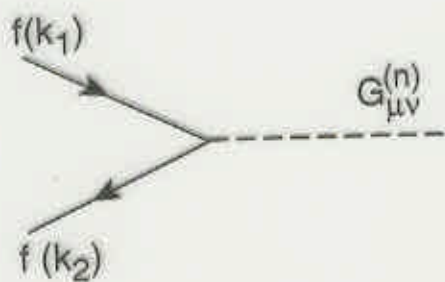
$$(\delta^2_{\mu} + \delta^2_a) \Phi = \sum_n \left[ \delta^2_{\mu} \phi^{(n)}(x^M) - \left(\frac{n}{R}\right)^2 \phi^{(n)}(x^M) \right] e^{inx^a/R}$$

↑ mass term with  
 $m = n/R$

# Feynman Rules

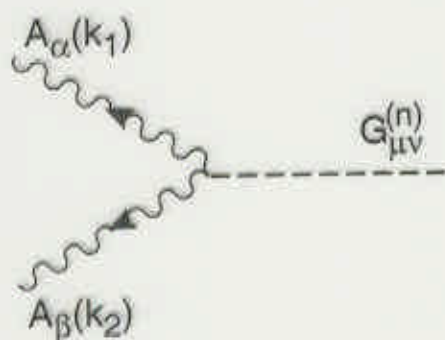
Giudice, Rattazzi,  
Wells

Han, Lykken,  
Zhang



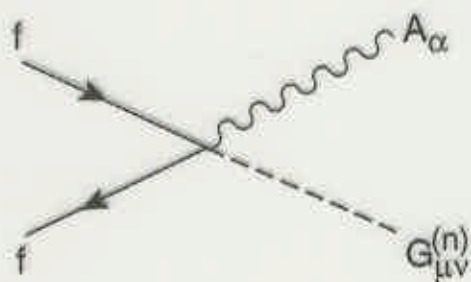
$$-\frac{i}{4\bar{M}_P} [W_{\mu\nu}^{(f)} + W_{\nu\mu}^{(f)}]$$

$$W_{\mu\nu}^{(f)} = (k_1 + k_2)_\mu \gamma_\nu \quad (53)$$



$$-\frac{i}{\bar{M}_P} [W_{\mu\nu\alpha\beta}^{(\gamma)} + W_{\nu\mu\alpha\beta}^{(\gamma)}]$$

$$W_{\mu\nu\alpha\beta}^{(\gamma)} = \frac{1}{2} \eta_{\mu\nu} (k_{1\beta} k_{2\alpha} - k_1 \cdot k_2 \eta_{\alpha\beta}) + \eta_{\alpha\beta} k_{1\mu} k_{2\nu} + \eta_{\mu\alpha} (k_1 \cdot k_2 \eta_{\nu\beta} - k_{1\beta} k_{2\nu}) - \eta_{\mu\beta} k_{1\nu} k_{2\alpha} \quad (54)$$



$$-\frac{i}{2\bar{M}_P} eQ (X_{\mu\nu\alpha} + X_{\nu\mu\alpha})$$

$$X_{\mu\nu\alpha} = \gamma_\mu \eta_{\nu\alpha} \quad (55)$$

Massless 0-mode  $\oplus$  all tower gravitons have same couplings to matter

## Two Classes of Collider Tests

### • Graviton Tower Emission

$$e^+e^- \rightarrow \gamma + G_n$$

$$q\bar{q} \rightarrow g + G_n$$

$$Z \rightarrow f\bar{f} + G_n$$

Giudice, Rattazzi, Wells  
Han, Lykken, Zhang  
Mirabelli, Perelstein, Peskin

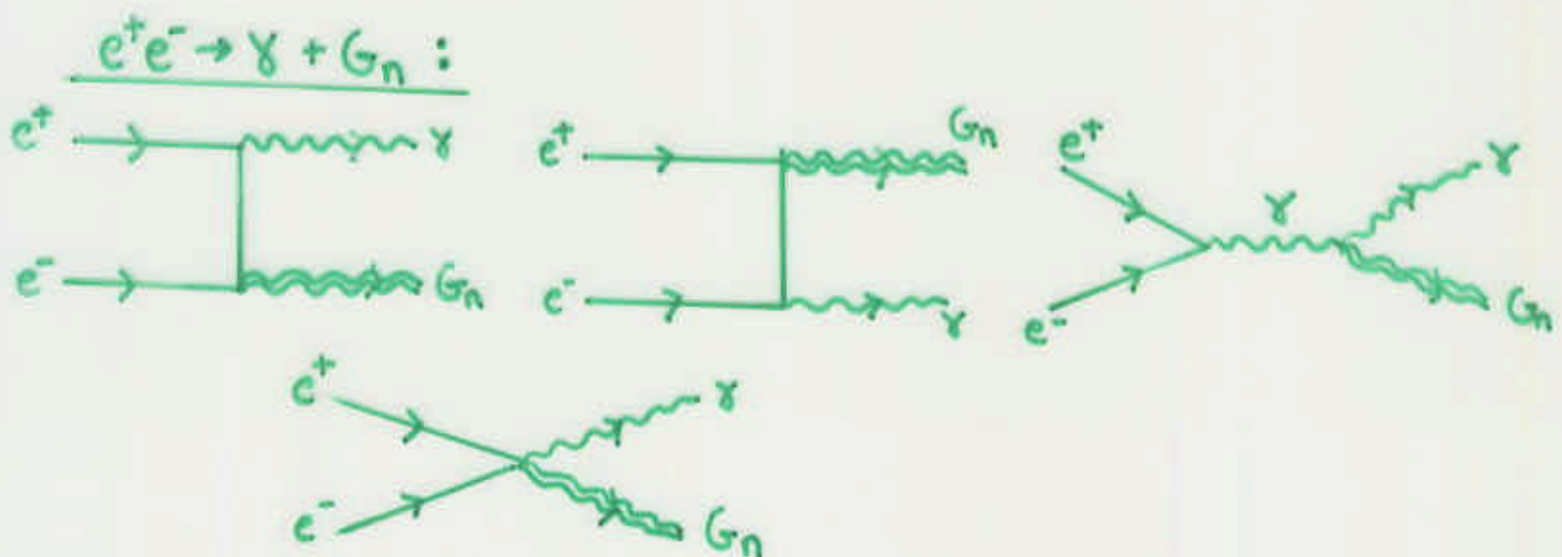
$G_n$  appears as  $\mathcal{L}_T$

Model Independent - Probes  $M_*$  directly

Sensitive to  $n$

Parameterized by effective density of states

$$\frac{d^2\sigma}{dm d\cos\theta} = \rho(m) \frac{d\sigma}{d\cos\theta} \quad \text{with} \quad \rho(m) = R^n \Omega_{n-1} m^{n-1}$$



## Graviton Emission : $e^+e^- \rightarrow \gamma + G^{(n)}$

Giudice,  
Ratazzi,  
Wells

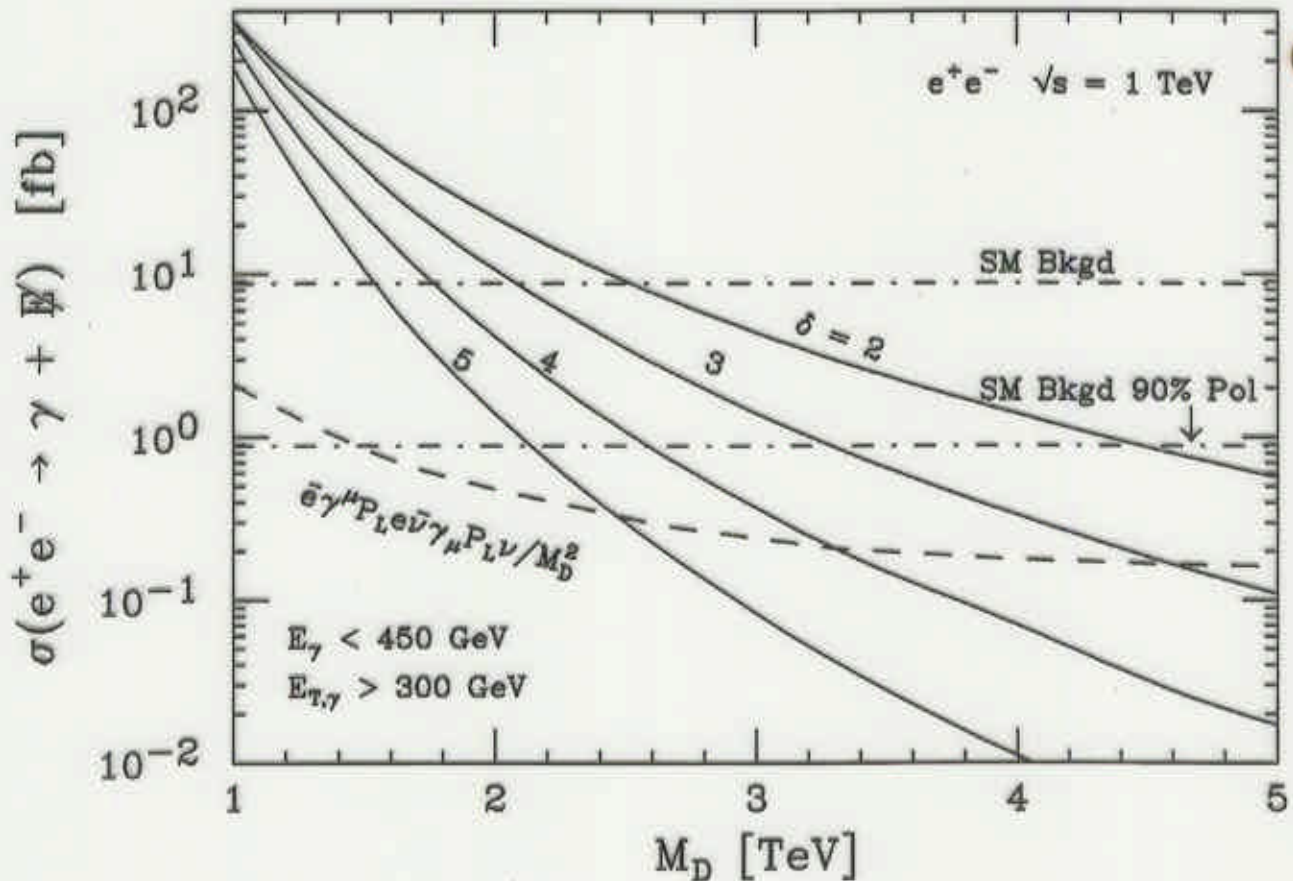


Figure 2: Total  $e^+e^- \rightarrow \gamma + \text{nothing}$  cross-section at a 1 TeV centre-of-mass energy  $e^+e^-$  collider. The signal from graviton production is presented as solid lines for various numbers of extra dimension ( $\delta = 2, 3, 4, 5$ ). The Standard Model background for unpolarized beams is given by the upper dash-dotted line, and the background with 90% polarization is given by the lower dash-dotted line. The signal and background are computed with the requirement  $E_\gamma < 450$  GeV in order to eliminate the  $\gamma Z \rightarrow \gamma \bar{\nu} \nu$  contribution to the background. The dashed line is the Standard Model background subtracted signal from a representative dimension-6 operator.

Note: signal  $\uparrow$  w/  $\sqrt{s} \uparrow$   
background  $\downarrow$  w/  $\sqrt{s} \uparrow$

$$pp \rightarrow g + G^{(n)}$$

ATLAS Simulation  
Vacavant, Hinchliffe

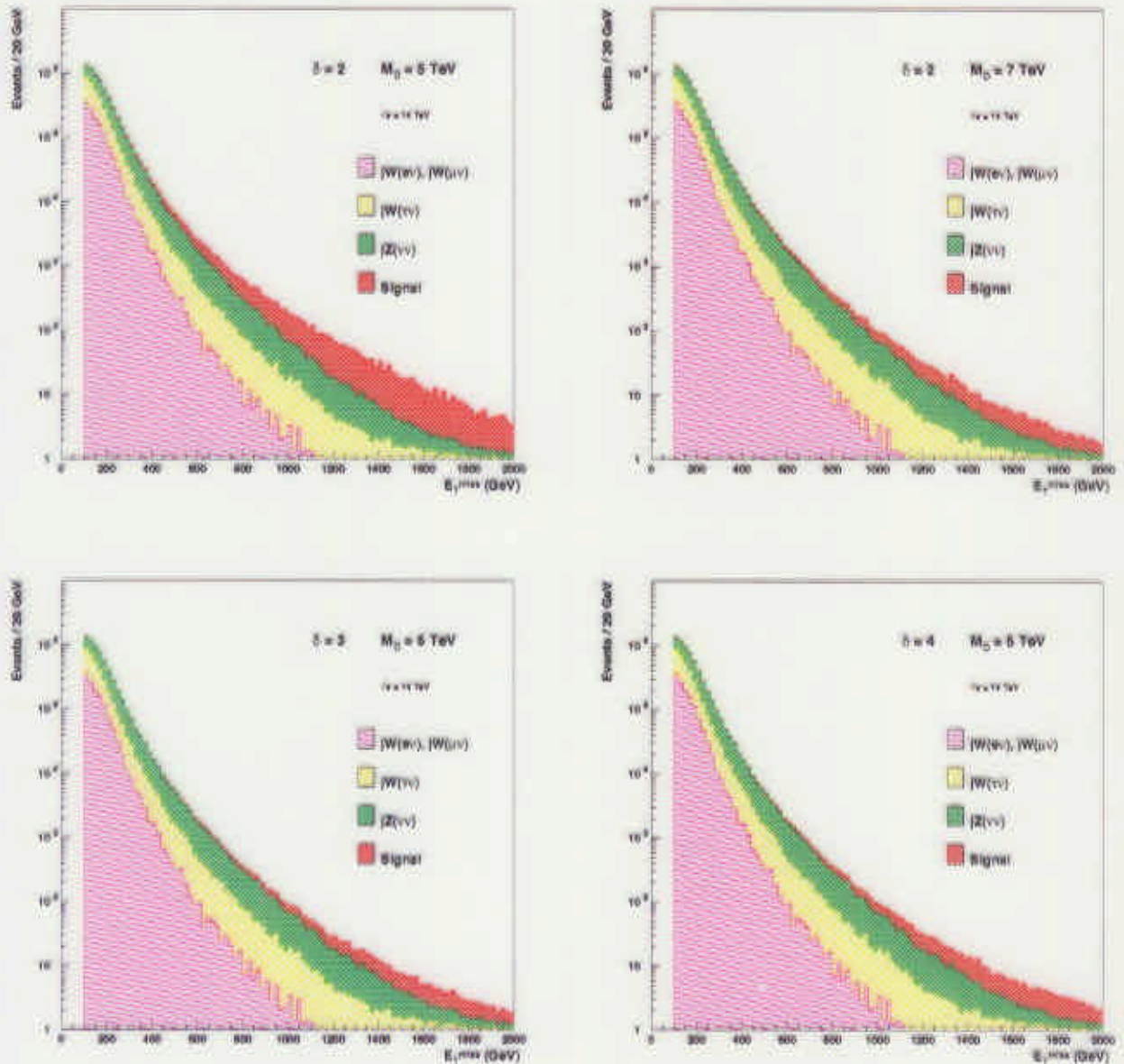


Figure 11: Distributions of the missing transverse energy in signal and in background events after the selection and for  $100 \text{ fb}^{-1}$  of integrated luminosity. Various cases  $(\delta, M_D)$  for the signal are shown.

Search Reach:

$n = 2$	$M_* = 4 - 7.5 \text{ TeV}$
3	4.5 - 5.9
4	5.0 - 5.3

• Graviton Tower Exchange  $XX \rightarrow G_n \rightarrow YY$

Search for 1) Deviations in SM processes  
2) New processes!

Graviton couples 1) universally to everything  
2) via  $T_{\mu\nu}$

Angular distributions reveal massive spin-2 exchange

Poor theoretical control!

$\Sigma$  KK propagators is divergent

$\Rightarrow$  Sensitivity to unknown ultraviolet physics

Approaches:

- 'Naive' cut-off JLH; Giudice et al; Han et al
- Brane fluctuations Bando et al
- Weakly Coupled String Theory Dudas, Mourad  
Accomando et al  
Cullen et al



## Effective Theory - Cut-off Approach

- Gravity becomes strong at  $M_*$   $\Rightarrow$  need full theory!
- Work in weak-field limit  $s \ll M_*^2$

Examine leading dimension-8 operators

$\Rightarrow$  contact interaction limit for  $G_n$  exchange

$\Rightarrow$  Constrain  $M_* |\lambda|^{-1/4}$

$$M = \frac{\lambda}{M_*^4} \left\{ \begin{array}{l} \bar{f} \gamma_\mu f \bar{l} \gamma^\mu l (p_f - p_{\bar{f}}) \cdot (p_l - p_{\bar{l}}) \\ + \bar{f} \gamma_\mu f \bar{l} \gamma_\nu l (p_f - p_{\bar{f}})^\nu (p_l - p_{\bar{l}})^\mu \end{array} \right\}$$

$$\frac{d\sigma}{dz} \sim A(1+z^2) + Bz$$

$$z = \cos \theta$$

$$- \frac{\lambda s^2}{M_*^4} [Cz^3 - D(1-3z^2)]$$

$$+ \frac{\lambda^2 s^4}{M_*^8} [E(1-3z^2+4z^4)]$$

$\Rightarrow$  Unique signal for spin-2 exchange!

# Angular Distributions for $e^+e^- \rightarrow f\bar{f}$

$\sqrt{s} = 500 \text{ GeV}$

$M_s = 1.5 \text{ TeV}$

Events  
bin

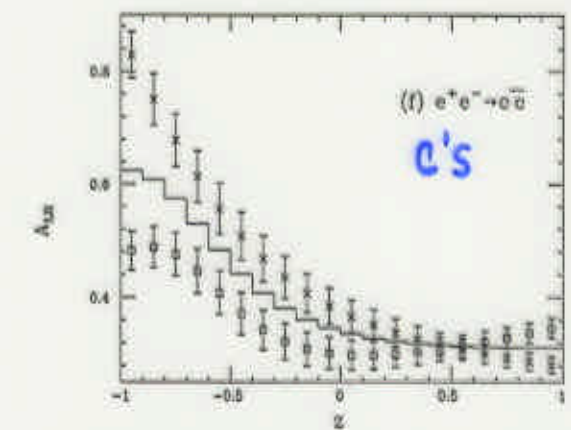
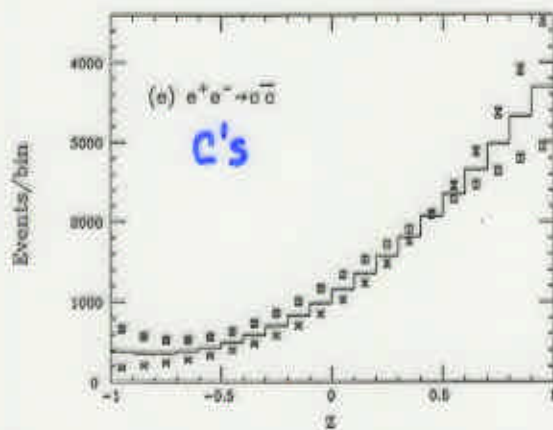
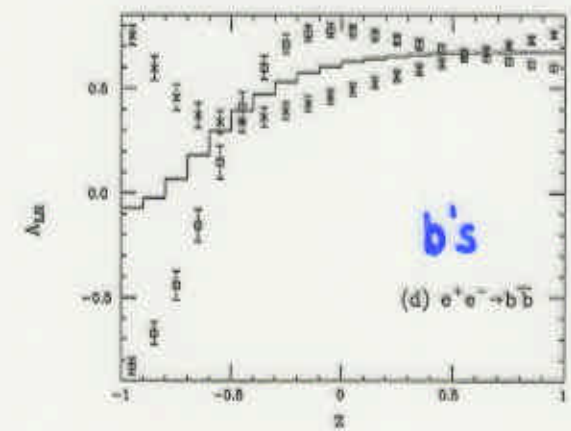
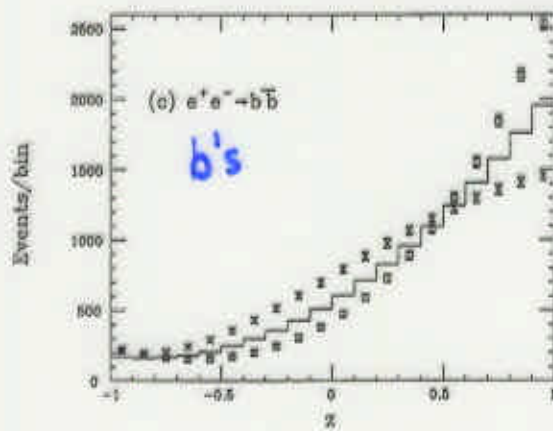
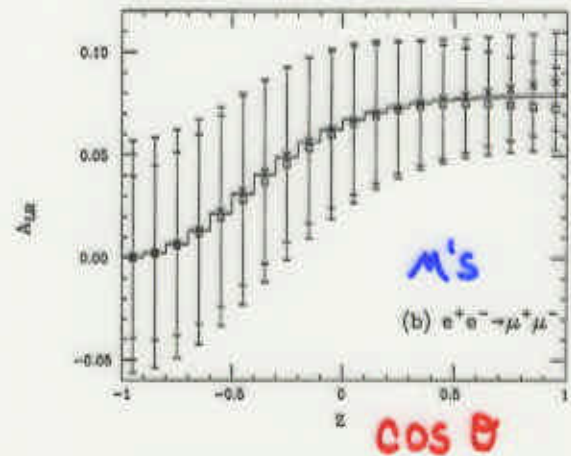
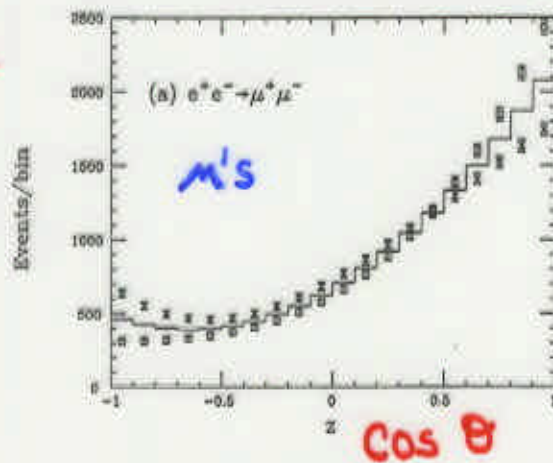
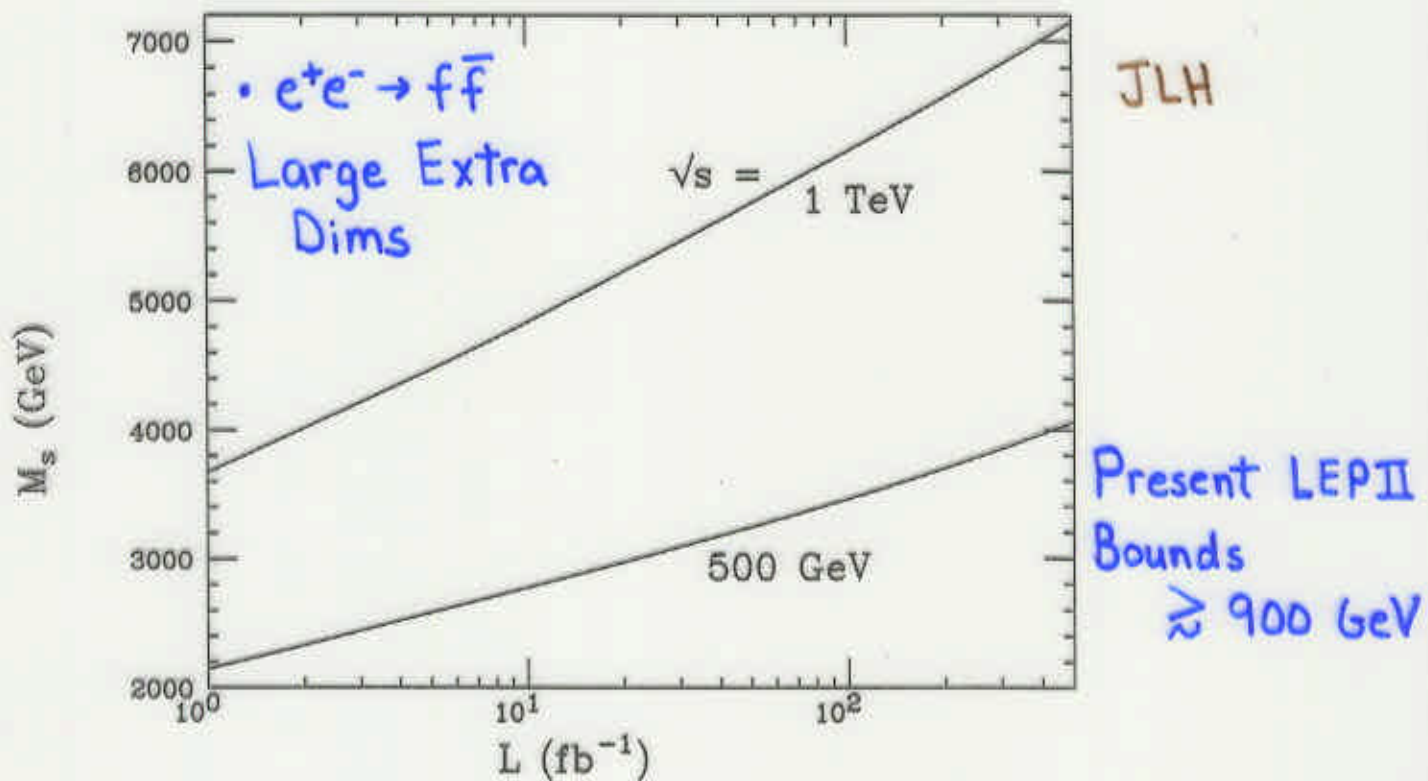
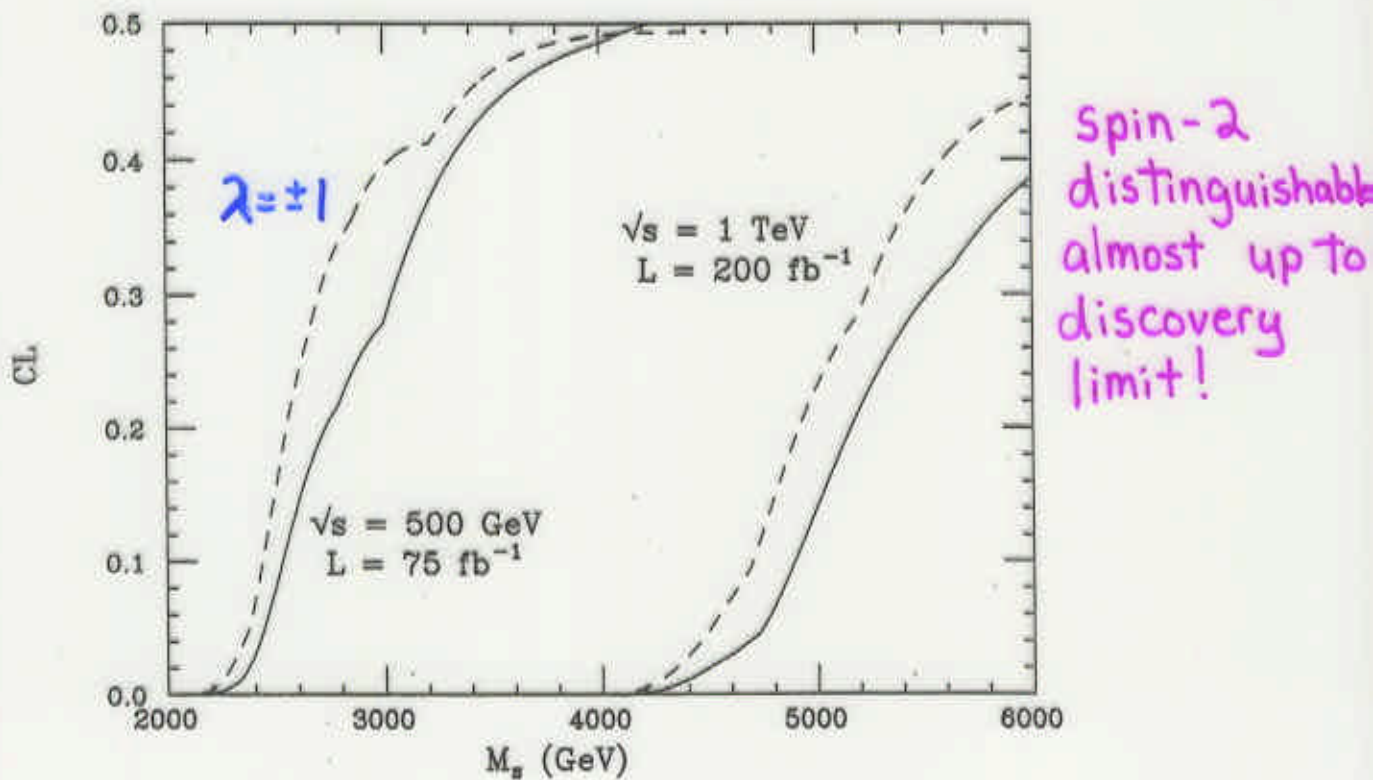


Figure 1: Bin integrated angular distribution and  $z$ -dependent Left-Right asymmetry for  $e^+e^- \rightarrow \mu^+\mu^-, b\bar{b}, c\bar{c}$ . In each case, the solid histogram represents the SM, while the 'data' points are for  $M_s = 1.5 \text{ TeV}$  with  $\lambda = \pm 1$ . The error bars correspond to the statistics in each bin.

# 95% CL Search Reach for Graviton Exchange



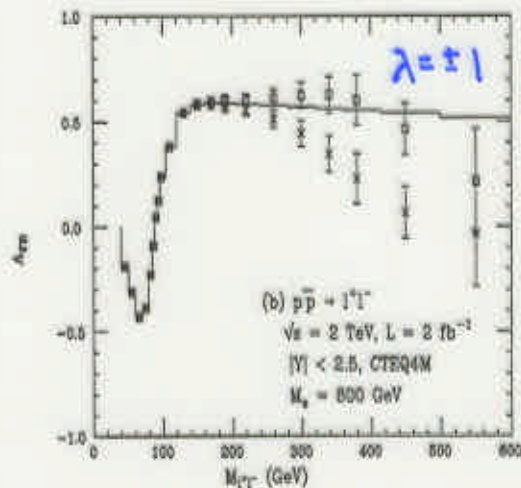
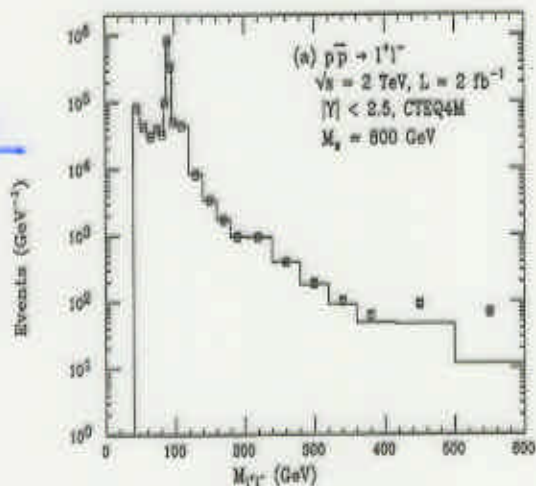
## Confidence Level of fit of spin-2 data to spin-1 hypothesis



Drell-Yan Production:  $q\bar{q} \rightarrow \gamma, Z, G^{(n)} \rightarrow l^+l^-$   
 $gg \rightarrow G^{(n)} \rightarrow l^+l^-$

Tevatron

$M_s = 800$   
GeV



JLH

LHC

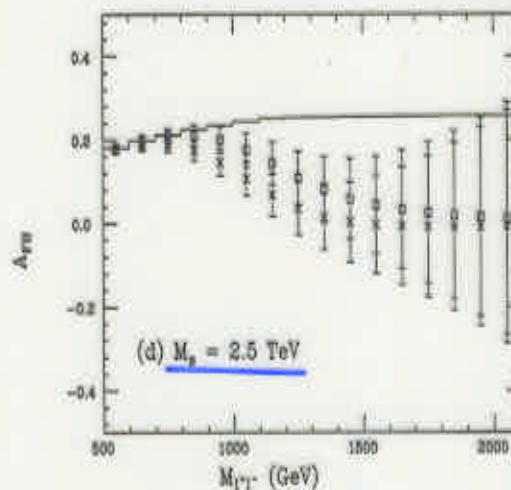
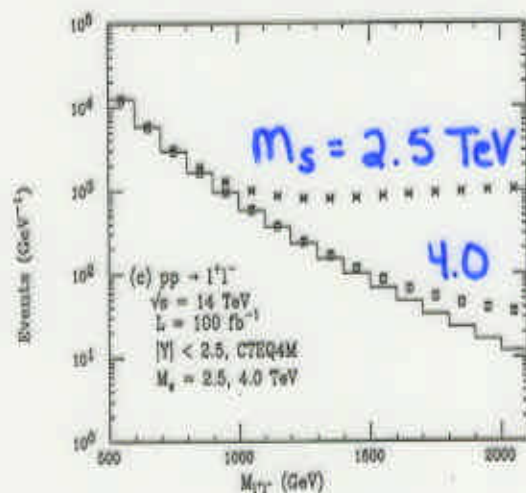


Figure 3: Bin integrated lepton pair invariant mass distribution and forward-backward asymmetry for Drell-Yan production at the Main Injector and the LHC. The SM is represented by the solid histogram. The data points represent graviton exchanges with (a)  $M_s = 800$  GeV and  $\lambda = +1$  or  $-1$ , (b)  $M_s = 800$  GeV and  $\lambda = +1$  and  $-1$ , (c)  $M_s = 2.5$  and  $4.0$  TeV and  $\lambda = +1$  or  $-1$ , (d)  $M_s = 2.5$  TeV and  $\lambda = +1$  and  $-1$ .

## Search Reach

Tevatron:  $M_s \sim 1-1.5$  TeV<sub>14</sub>

LHC:  $M_s \sim 4-5.2$  TeV

## 5-D Standard Model

(Antoniadis et al)

Do all  $n$  dimensions have to be the same size?

$$\text{Let } R^n = R_1^p R_2^{n-p} \quad \text{with } R_1 \sim \text{large} \\ R_2 \sim \text{small} \sim 1/\text{TeV} \sim 1/m_*$$

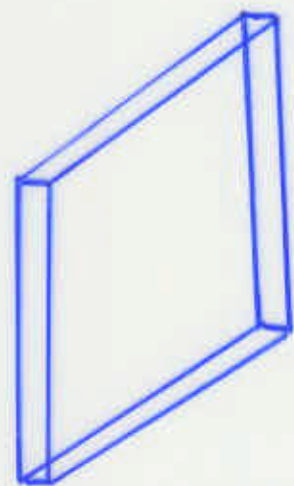
$$\Rightarrow m_{p1}^2 = R_1^p m_*^{p-n} m_*^{n+2} \\ = R_1^p m_*^{p+2} \quad \text{with } 2 \leq p \leq 6$$

SM fields can propagate in small  $R_2^{n-p}$  dimensions!

5-DSM: 1 extra small dimension

Degenerate  $\gamma/W/Z/g$  KK towers  
with  $g^{(n)} = \sqrt{2} g^{(0)}$

Precision EW data:  $m_c \gtrsim 3-4 \text{ TeV}$   
(Rizzo, Wells)



Fat-branes

Modifies  $\Sigma_T$  signatures  
by form factor  $\Rightarrow$  reduced signal  
(De Rújula et al)

$\sim \lambda \sim \text{TeV}$

## Search Reach:

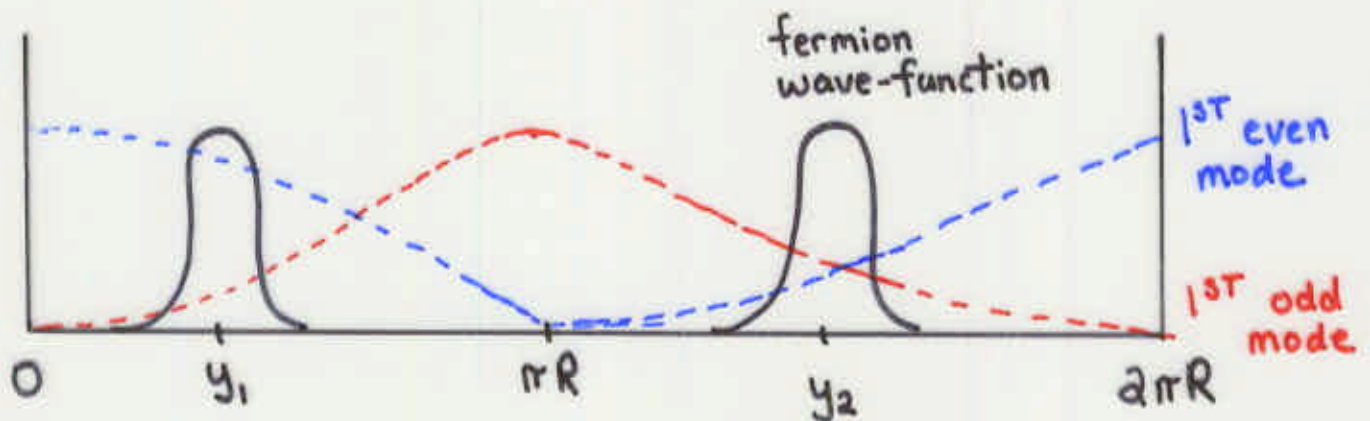
### Experiment

### $m_c$ Reach (TeV)

LEP II		3.1
Tevatron	$2 \text{ fb}^{-1}$	1.1
	$20 \text{ fb}^{-1}$	1.3
LHC	$100 \text{ fb}^{-1}$	6.3
NLC	$\sqrt{s} = 0.5 \text{ TeV}$	13
	$1.0 \text{ TeV}$	23
	$1.5 \text{ TeV}$	31

## Separated Fermions

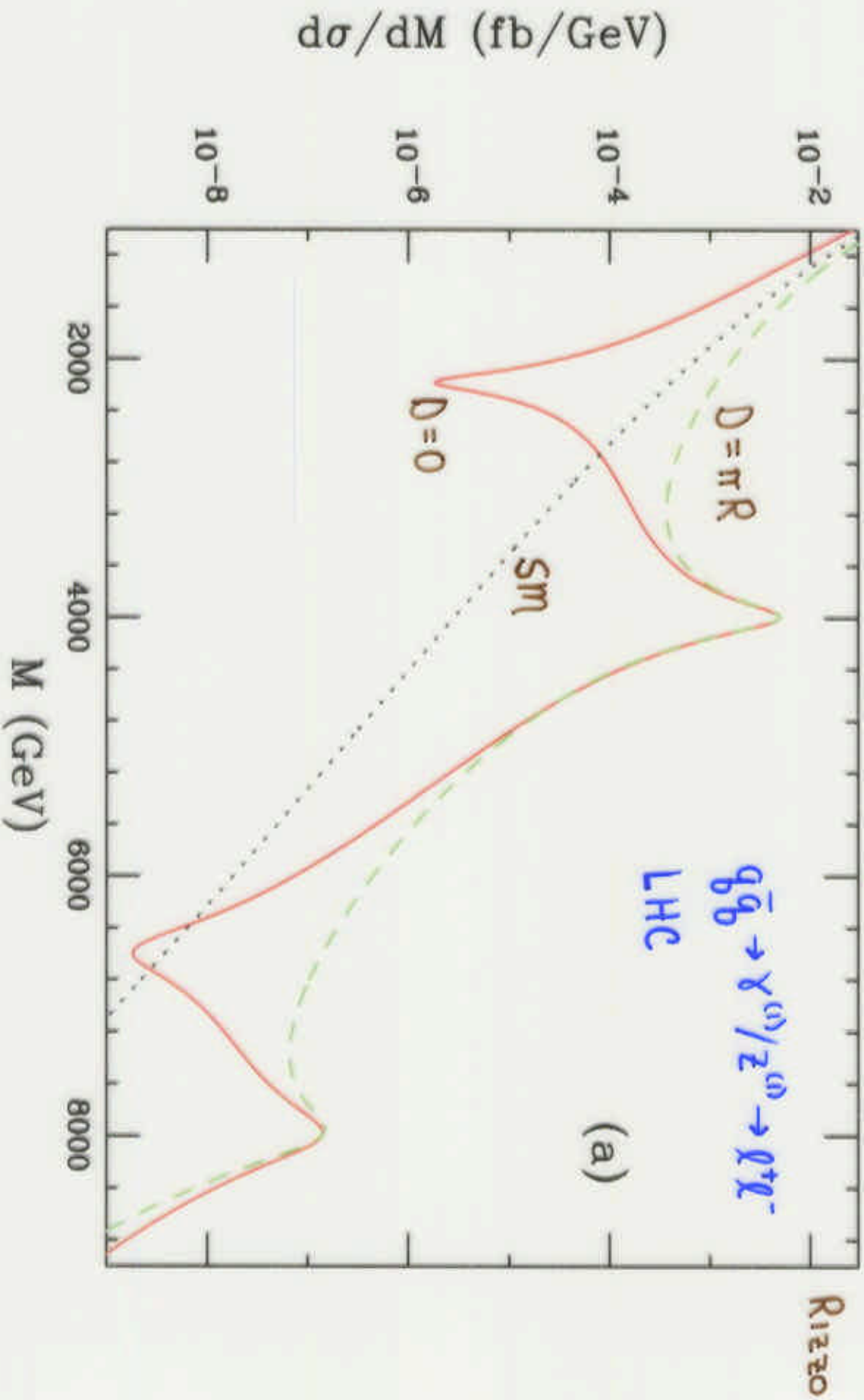
(Arkani-Hamed, Schmaltz)



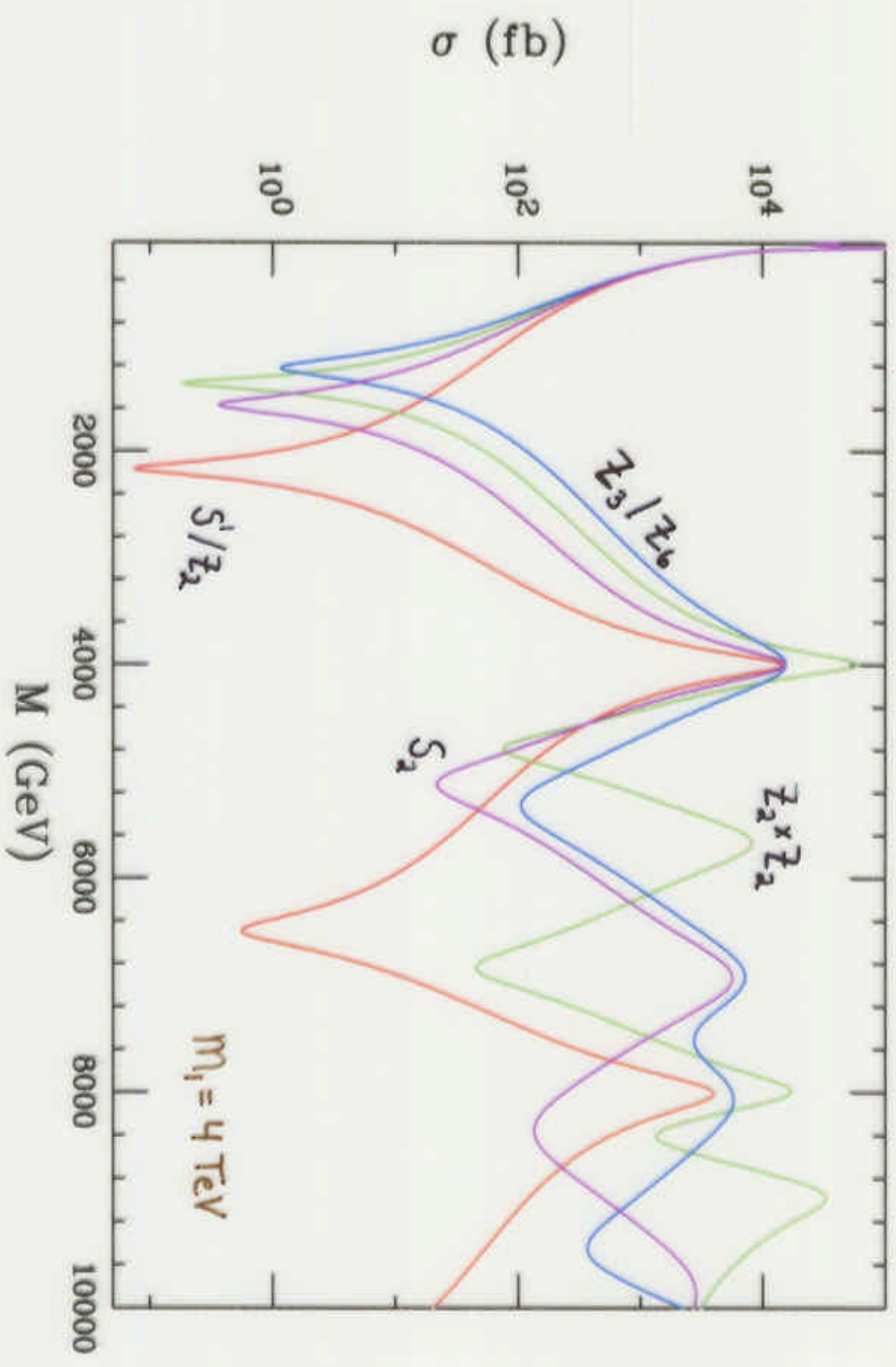
KK gauge coupling to fermions:

$$= \sqrt{2} g^{(0)} \int dy \bar{\psi}_1(y) \psi_2(y) G(y)$$

$D =$  separation of fermions in 5<sup>th</sup> dimension



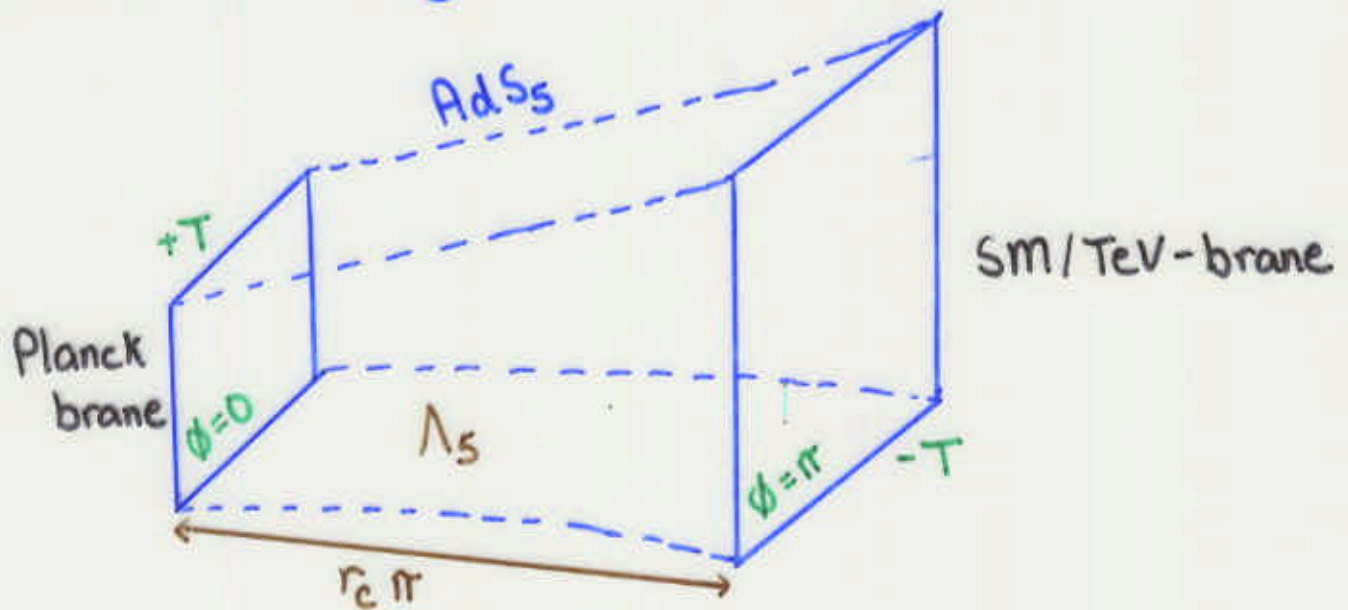
KK Excitation Pattern  $\mu^+\mu^- \rightarrow e^+e^-$



Can determine model of compactification and # of dimensions!



# Localized Gravity ala Randall-Sundrum



Bulk = Slice of  $AdS_5$

Two 3-branes at  $S_1/\mathbb{Z}_2$  orbifold fixed points

5-D, non-factorizable geometry

Solutions to Einstein's Eqn [w/ 4-D Poincare']

$$ds^2 = e^{-2kr_c|\phi|} \eta_{\mu\nu} dx^\mu dx^\nu + r_c^2 d\phi^2$$

Warp factor

$0 \leq |\phi| \leq \pi$   
 $r_c =$  compactification radius

where  $\Lambda_5 = -24 m_s^3 k^2$

## 4-D Effective Action:

$$\bar{m}_{p1}^2 = \frac{m_5^3}{k} (1 - e^{-2kr_c\pi}) \Rightarrow k \sim m_5 \sim \bar{m}_{p1}$$

no additional hierarchies!

## Physical scales:

$$\Lambda_\emptyset = e^{-kr_c|\emptyset|} \bar{m}_{p1}$$

For TeV-brane at  $\emptyset = \pi$

$$\Lambda_\pi = e^{-kr_c\pi} \bar{m}_{p1} \sim \text{TeV} \quad \text{if } kr_c \sim 11-12$$

$\Rightarrow$  hierarchy generated by an exponential!

stabilized via  
Goldberger + Wise

## Phenomenology governed by

$k/\bar{m}_{p1}$  and  $\Lambda_\pi \Rightarrow$  only 2 free parameters

## 5-D curvature:

$$|R_5| = 20k^2 < m_5^2$$

suggests  $k/\bar{m}_{p1} < 0.1$

(neglecting  
higher-order  
curvature terms)

## 4-D Effective Theory:

Linear expansion of flat metric

$$G_{\alpha\beta} = e^{-2\kappa r_c |\vartheta|} (\eta_{\alpha\beta} + \kappa h_{\alpha\beta}) \quad \kappa = 2 m_5^{-3/2}$$

Expand into KK tower

$$h_{\alpha\beta}(x, \vartheta) = \sum_{n=0}^{\infty} h_{\alpha\beta}^{(n)}(x) \frac{\chi_h^{(n)}(\vartheta)}{\sqrt{r_c}}$$

Employ gauge  $\eta^{\alpha\beta} \partial_\alpha h_{\beta\gamma}^{(n)} = 0 + \eta^{\alpha\beta} h_{\alpha\beta}^{(n)} = 0$

Demand  $\int_{-\pi}^{\pi} d\vartheta e^{-2\kappa r_c |\vartheta|} \chi_h^{(m)} \chi_h^{(n)} = \delta_{mn}$

$$\chi_h^{(n)}(\vartheta) = \frac{e^{2\kappa r_c |\vartheta|}}{N_n} \left[ J_2\left(\frac{m_n}{\kappa} e^{\kappa r_c |\vartheta|}\right) + \alpha_n Y_2\left(\frac{m_n}{\kappa} e^{\kappa r_c |\vartheta|}\right) \right]$$

For TeV-brane:

$$\begin{aligned} m_n &= x_n \kappa e^{-\kappa r_c \pi} & \text{with } J_1(x_n) &= 0 \\ &= x_n \frac{\kappa}{\bar{m}_{pl}} \Lambda \pi \end{aligned}$$

$\Rightarrow$  KK excitations are not evenly spaced!

## Interactions

$$\mathcal{L} \sim \frac{1}{m_s^{3/2}} T^{\alpha\beta}(x) h_{\alpha\beta}(x, \varnothing = \pi)$$

$$= \frac{1}{\bar{m}_{\rho_1}} T^{\alpha\beta}(x) h_{\alpha\beta}^{(0)}(x) - \frac{1}{\Lambda_\pi} T^{\alpha\beta}(x) \sum_{n=1}^{\infty} h_{\alpha\beta}^{(n)}(x)$$

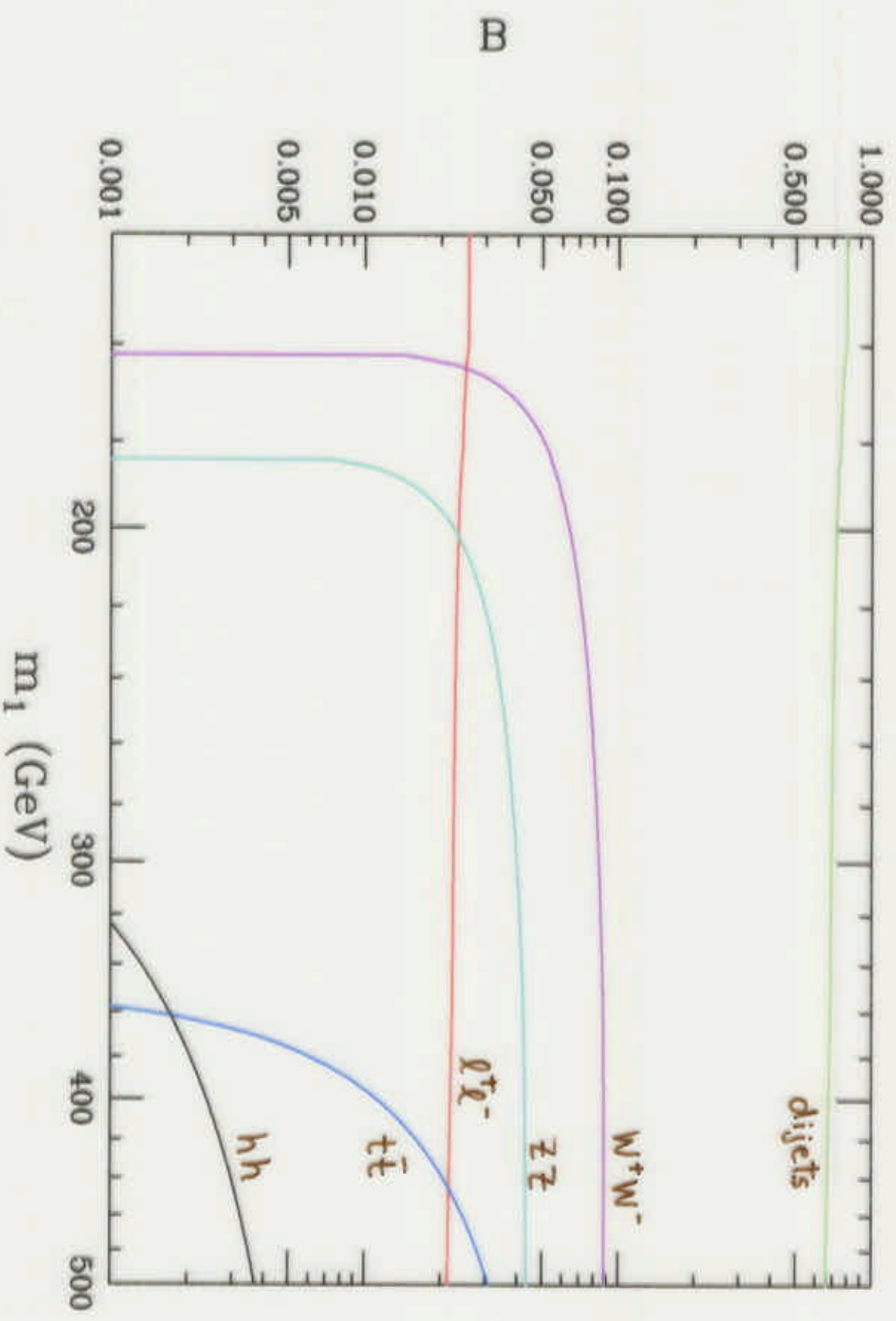
Zero-mode decouples

TeV-suppressed  
 $\Rightarrow$  can be directly produced!

## Phenomenology

- Graviton resonance production
- Graviton contributions to EW oblique parameters
- 'Light, skinny' Gravitons  $[k/\bar{m}_{\rho_1} \lesssim 0.01]$   
 Graviton emission
- Below resonance exchange  
 "Contact interaction" limits

# Graviton Branching Fractions



$$B_{\gamma\gamma} = 2 B_{l^+l^-}$$

# KK Graviton Drell-Yan Spectrum

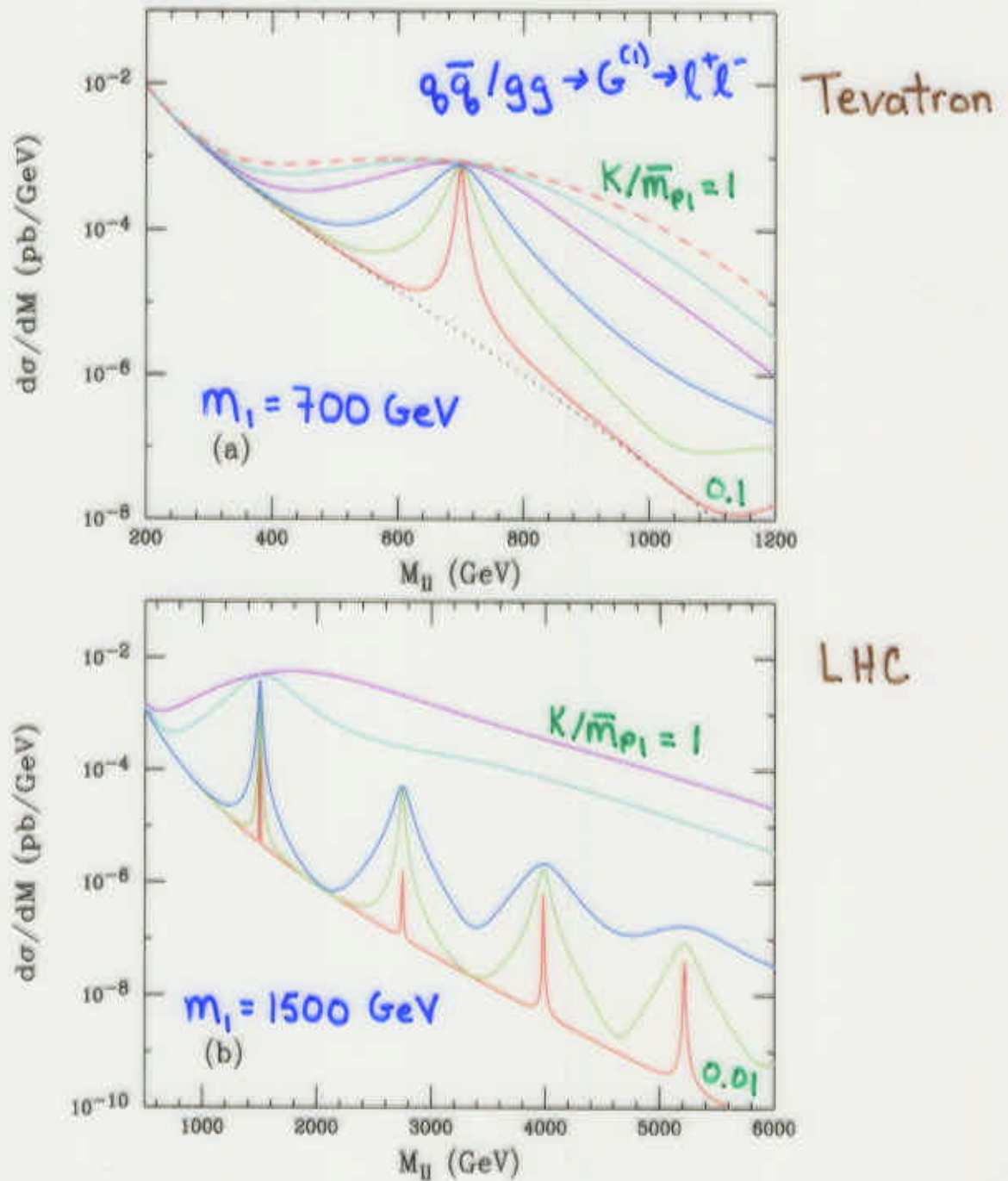
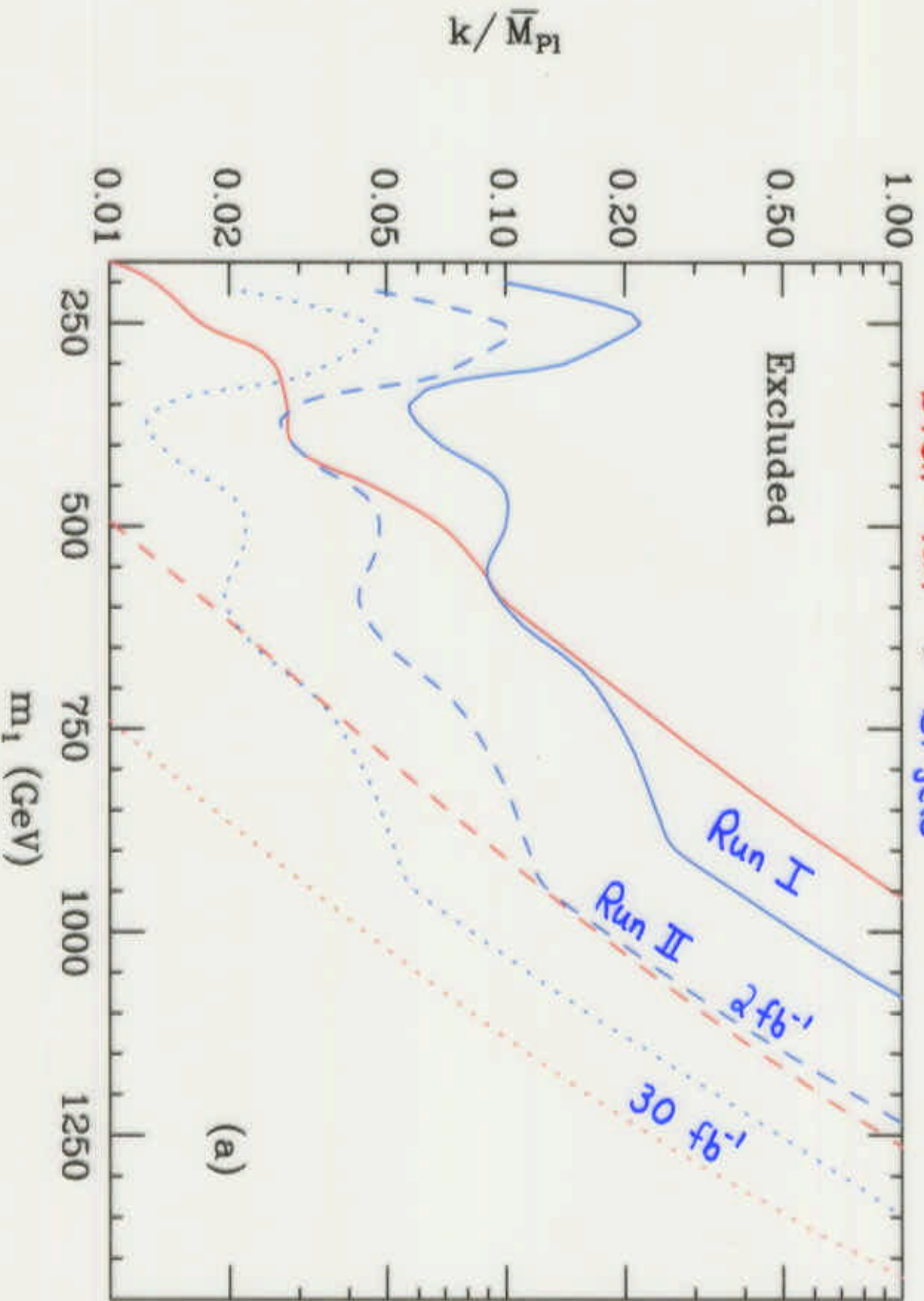
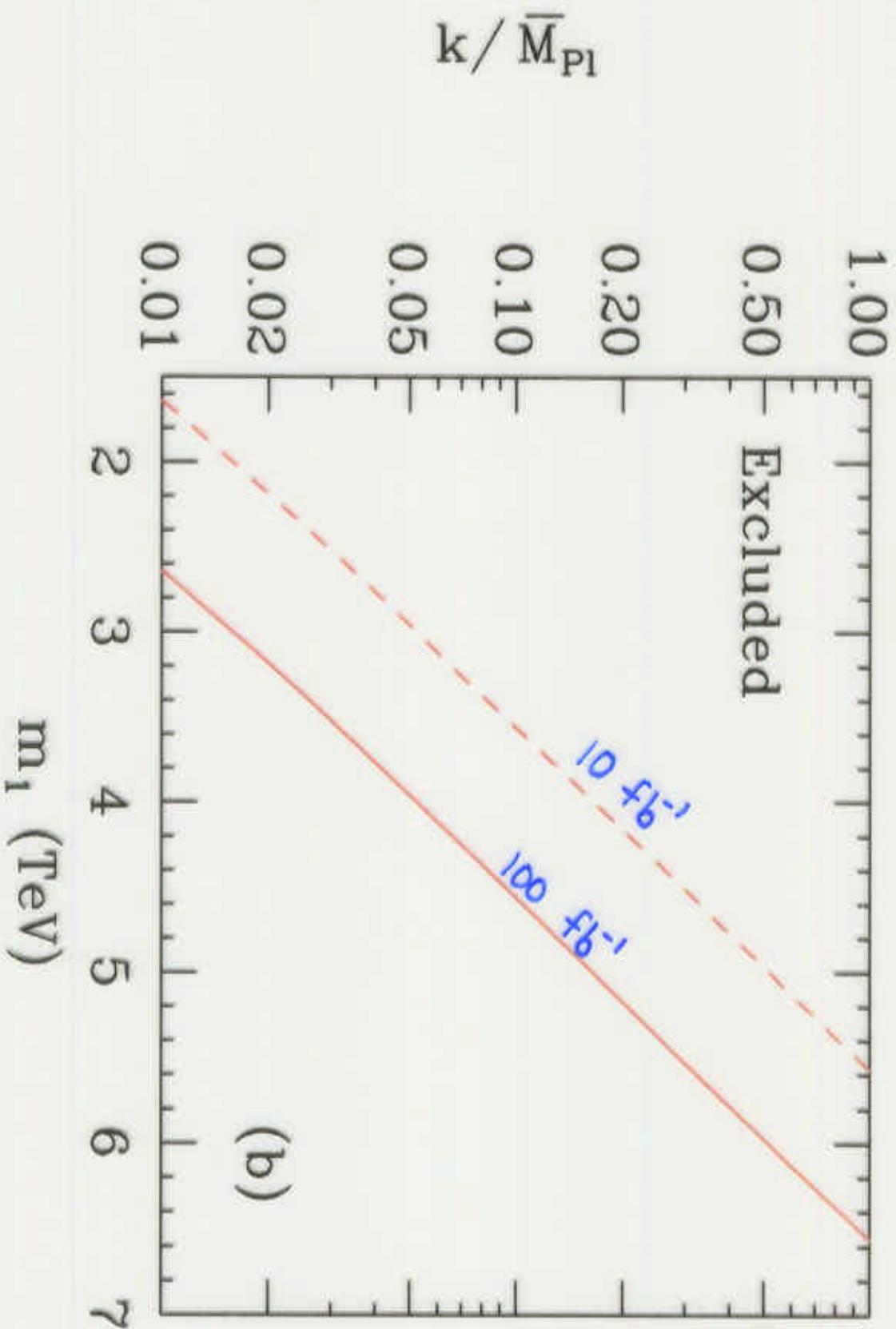


Figure 17: Drell-Yan production of a (a) 700 GeV KK graviton at the Tevatron with  $k/\bar{M}_{Pl} = 1, 0.7, 0.5, 0.3, 0.2,$  and  $0.1$ , respectively, from top to bottom; (b) 1500 GeV KK graviton and its subsequent tower states at the LHC. From top to bottom, the curves are for  $k/\bar{M}_{Pl} = 1, 0.5, 0.1, 0.05,$  and  $0.01$ , respectively.

## Direct Bump Searches - Tevatron

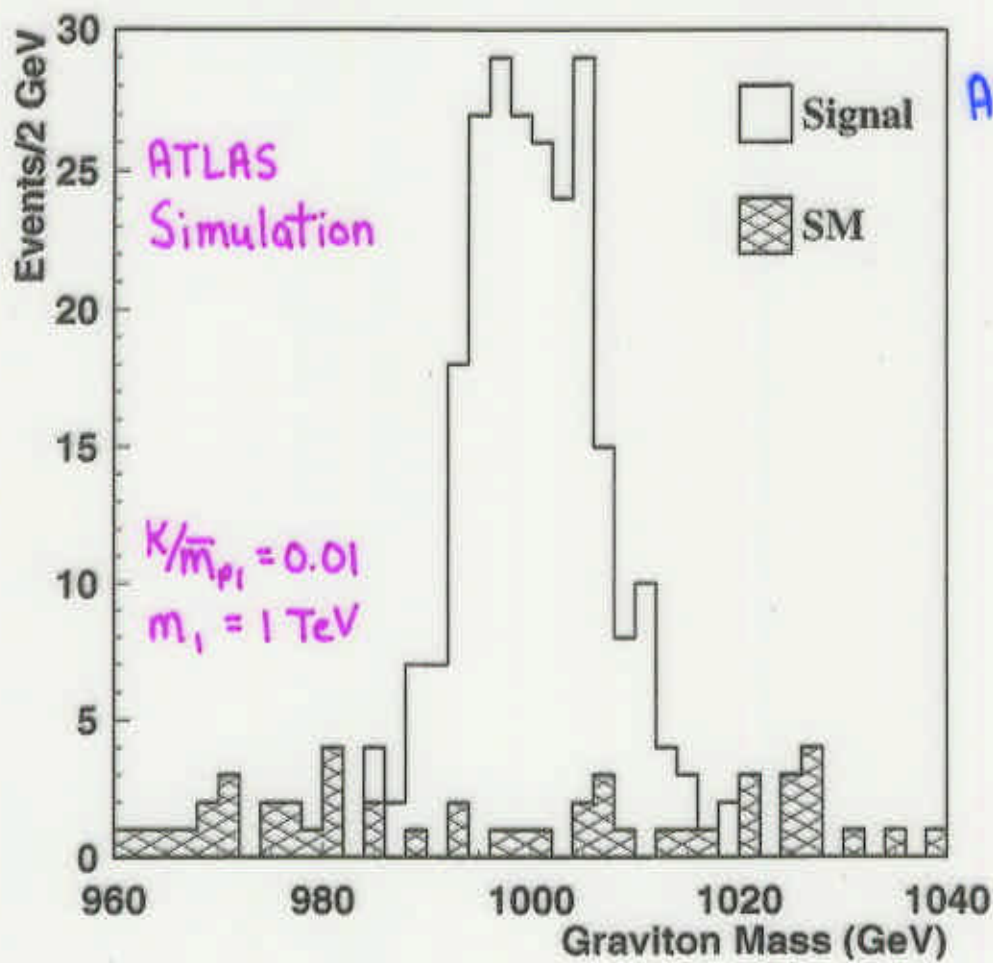
Drell-Yan  $\rightarrow$  Di-jets

Drell-Yan @ LHC





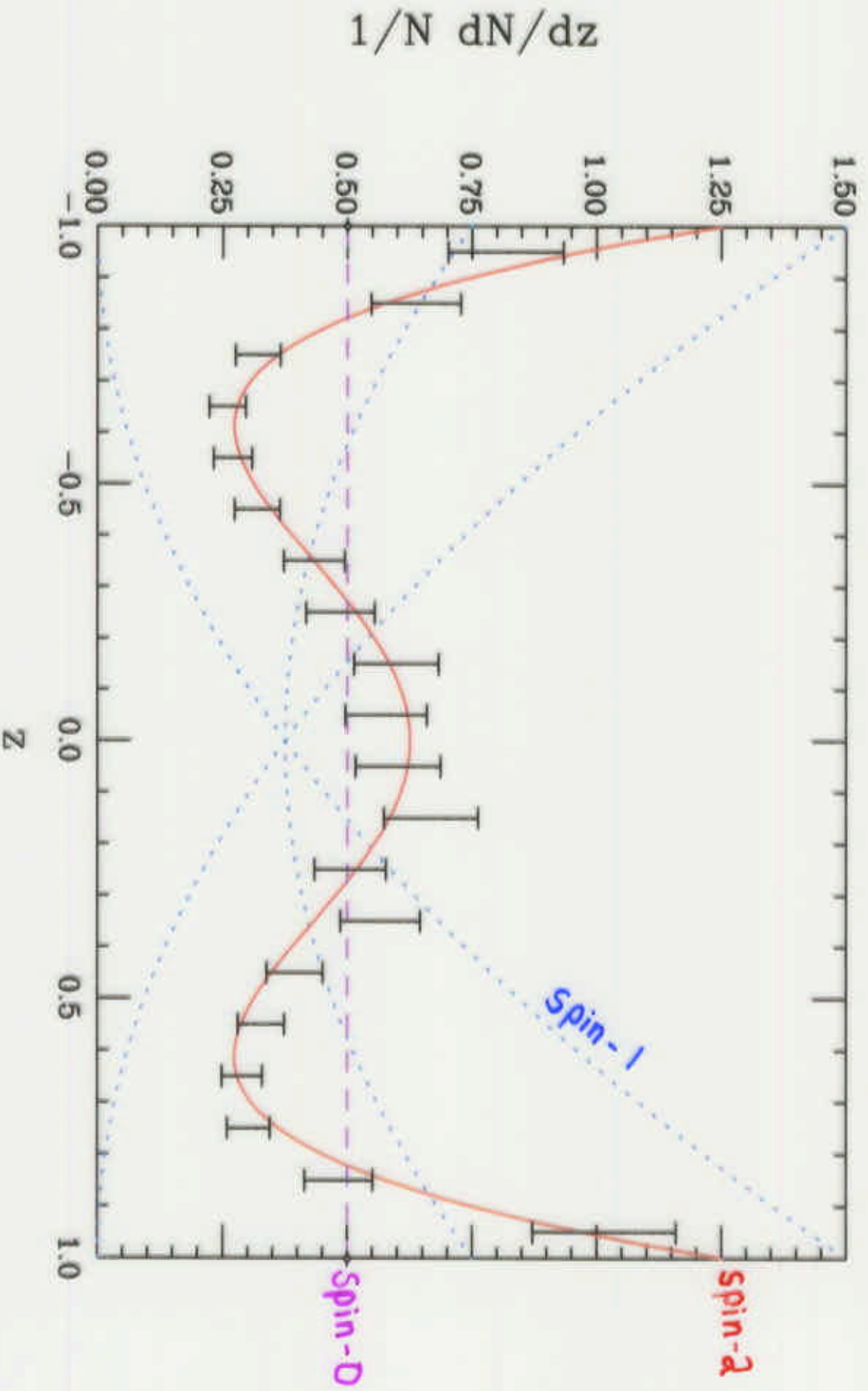
## Narrow-width Graviton Resonance



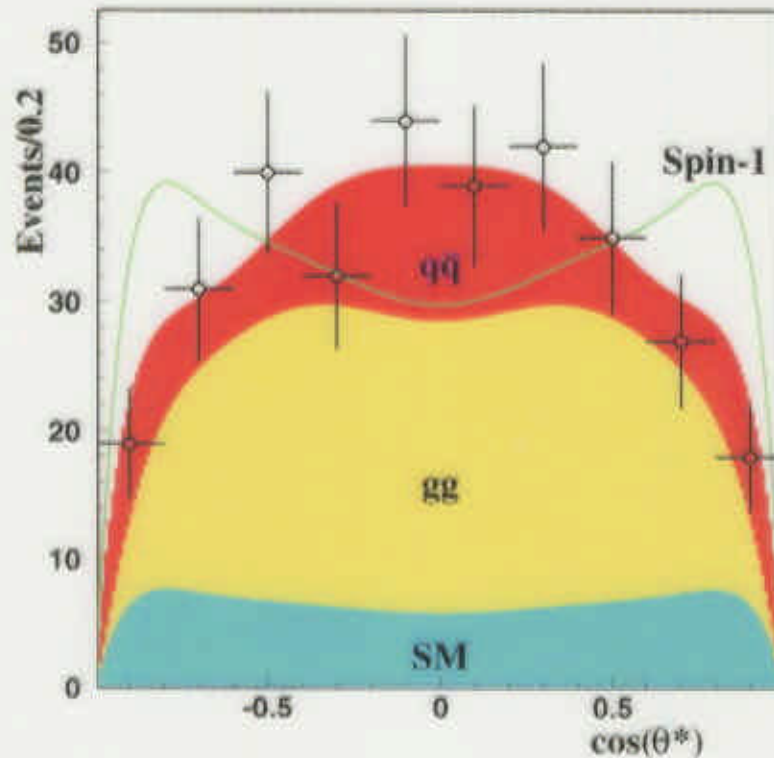
Allanach et al

ATLAS search reach:  $m_{\pi_1} \sim 1830 \text{ GeV}$  for  $K/\bar{m}_{\pi_1} = 0.01$

# On-Resonance Spin Determination



# Spin-2 Determination from Drell-Yan



Allanach et al  
ATLAS Simulation

$m_1 = 1 \text{ TeV}$   
 $100 \text{ fb}^{-1}$

**Figure 4:** The angular distribution of data (points with errors) in the test model for  $m_G = 1000 \text{ GeV}$  and  $100 \text{ fb}^{-1}$  of integrated luminosity. The stacked histograms show the contributions from the Standard Model (SM),  $gg$  production ( $gg$ ) and  $q\bar{q}$  production ( $q\bar{q}$ ). The curve shows the distribution expected from a spin-1 resonance.

distributions, defined as

$$L = x_q \cdot f_q(\theta^*) \cdot A_q(M, \theta^*)/I_q(M) + x_g \cdot f_g(\theta^*) \cdot A_g(M, \theta^*)/I_g(M) + x_{DY} \cdot f_{DY}(\theta^*) \cdot A_{DY}(M, \theta^*)/I_{DY}(M) \quad (4.1)$$

where  $x_i$  is the fraction of the events from each contributing process,  $f_i(\theta^*)$  is the angular distribution of the process,  $A_i(M, \theta^*)$  is the acceptance of the detector as a function of the mass of the electron pair and  $\theta^*$ , and

$$I_i(M) = \int_{-1}^1 f_i(\theta^*) \cdot A_i(M, \theta^*) d \cos \theta^* \quad (4.2)$$

$i = q, g, DY$  for the processes  $q\bar{q} \rightarrow G$ ,  $gg \rightarrow G$ , and  $q\bar{q} \rightarrow Z/\gamma^*$  respectively. Only the shape of the distribution is used in the statistical tests, and the coefficients  $x$  are constrained such that

$$x_q + x_g + x_{DY} = 1 \quad (4.3)$$

In order to evaluate the discovery reach of the experiment, in terms of its ability to reveal the spin-2 nature of the resonance, the following procedure was followed, intended to mimic an ensemble of possible experimental runs:

# Graviton Drell-Yan cross section

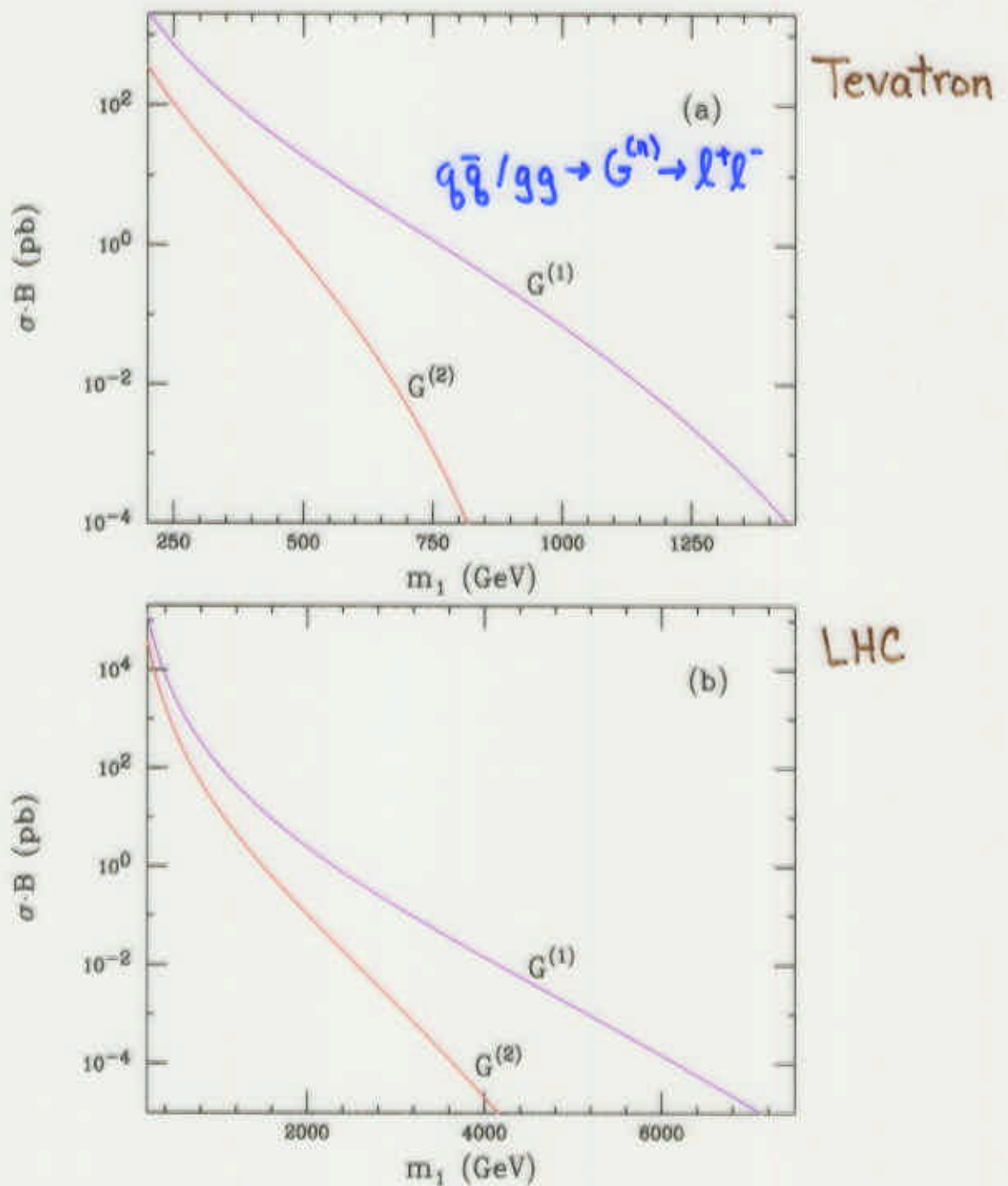


Figure 18: Cross sections for Drell-Yan production at the (a) Tevatron and (b) LHC of the first two graviton KK states coupling to the SM on the wall as a function of  $m_1$ . The upper (lower) curve in each case is for the first (second) KK state. Here, we have set  $k/\sqrt{M_{Pl}} = 0.1$ .

## Constraints from Precision Measurements

Corrections to Gauge Boson self-energies:



For large extra dimensions:  $M_* > 1.2 - 1.5 \text{ TeV}$

[Han, Marfatia, Zhang]

For RS, KK sum must be performed explicitly.

Gravity becomes strong for  $p > \Lambda_\pi$

$\Rightarrow$  introduce cut-off  $\Lambda_{\text{cut}} = \Lambda_\pi \lambda$

$$\Pi(p^2) = \frac{1}{48\pi^2} \sum_n \left[ \frac{\Lambda_c}{m_n} \right]^4 \left[ \frac{1}{3} + \frac{4m_n^2}{\Lambda_c^2} + \frac{10m_n^4}{\Lambda_c^4} + \frac{10m_n^6}{\Lambda_c^6} \ln \frac{m_n^2/\Lambda_c^2}{1 + \frac{m_n^2}{\Lambda_c^2}} \right]$$

Data constrains oblique parameters: global fit

$$S = -0.04 \pm 0.10$$

$$T = -0.06 \pm 0.11$$

# KK Graviton contributions to oblique parameters

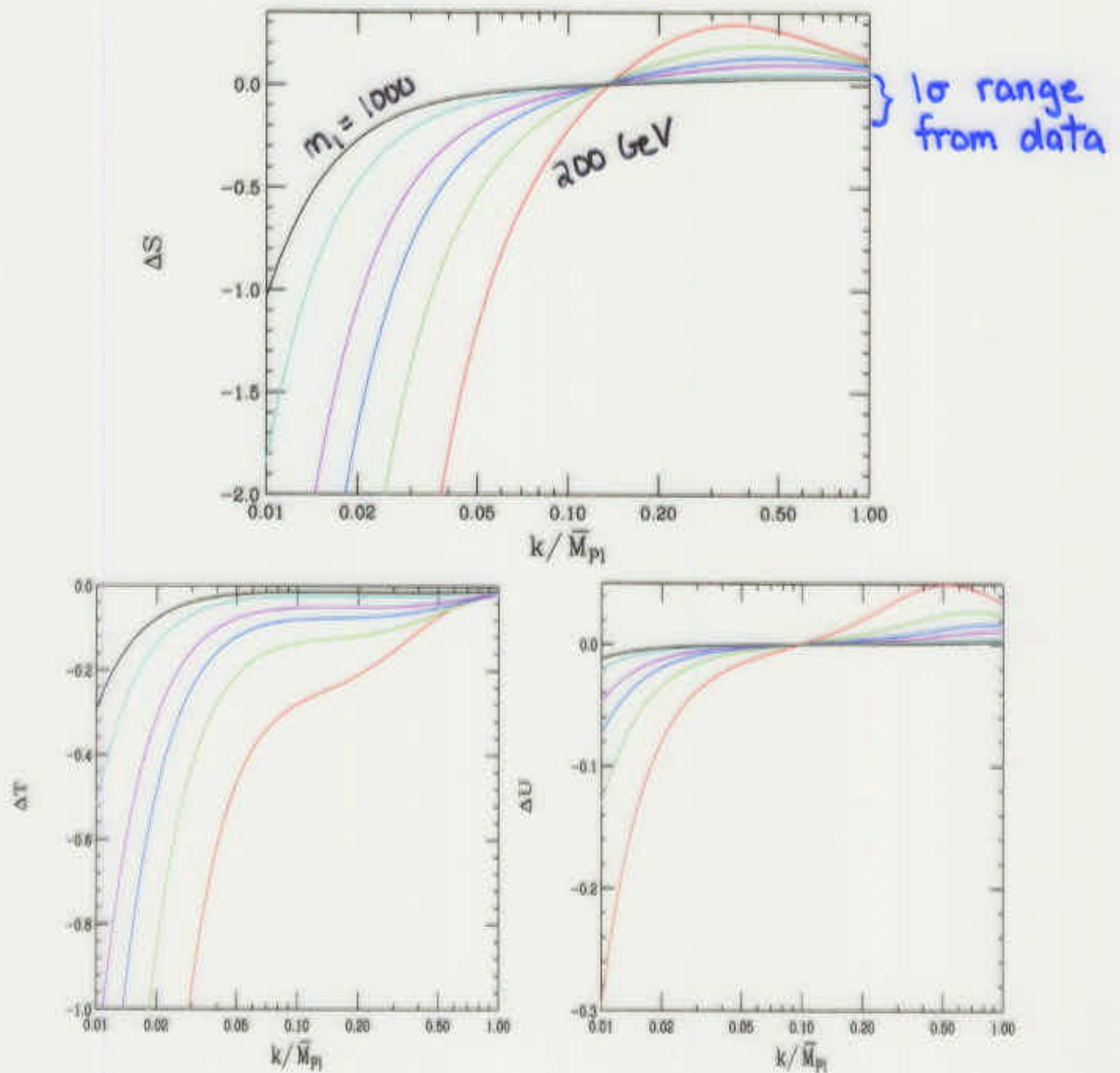
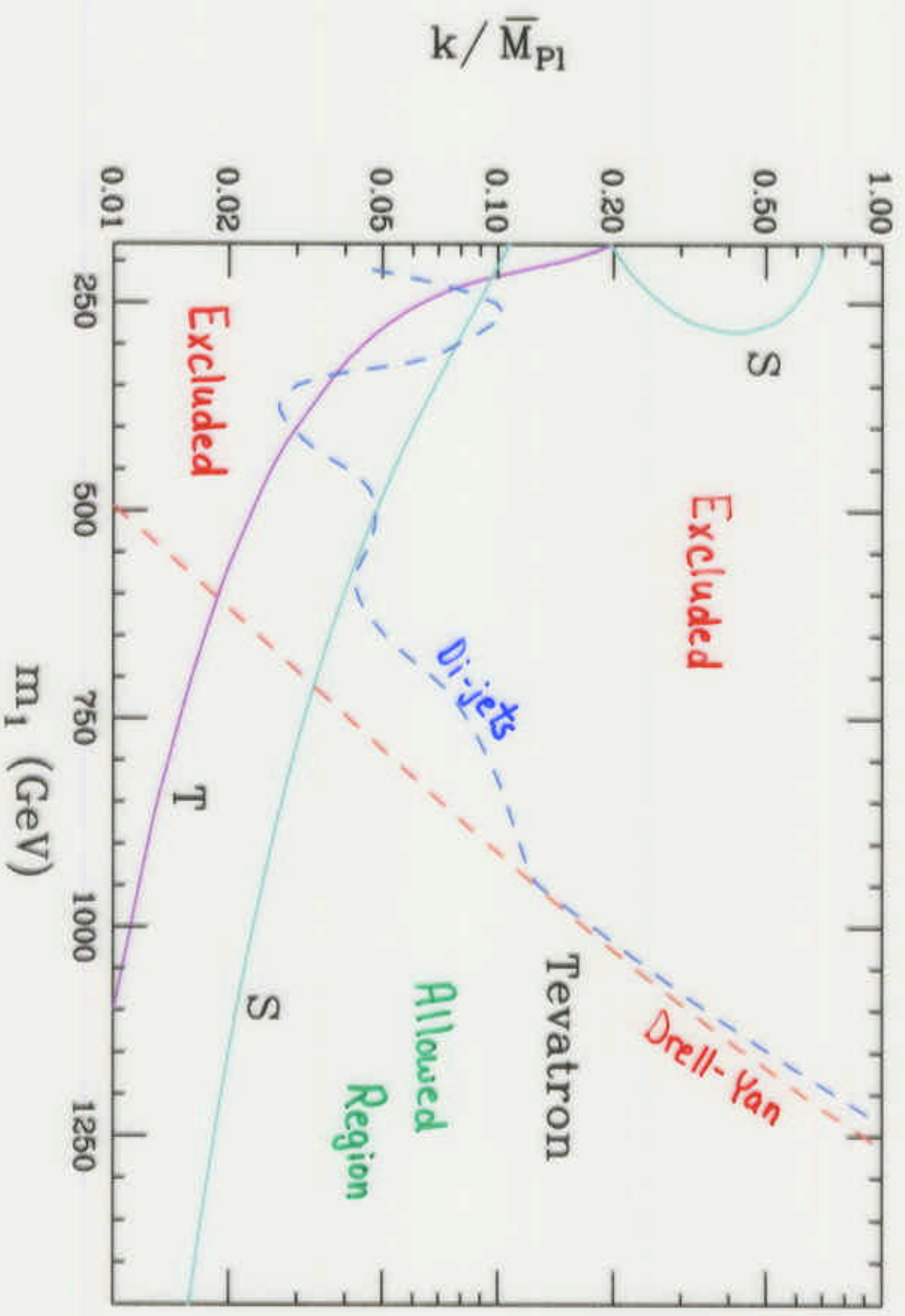
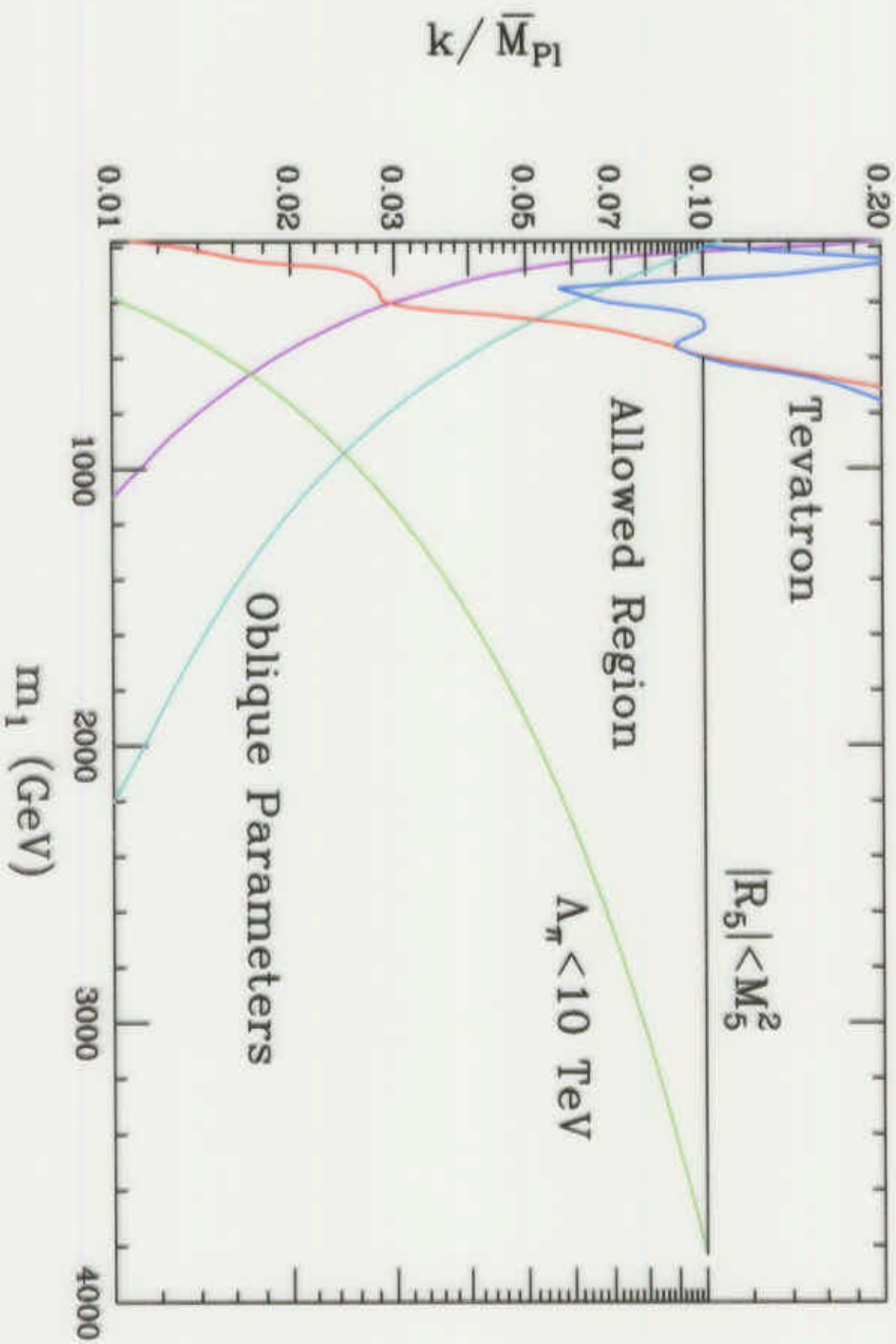


Figure 12: Shifts in the oblique parameters  $S$ ,  $T$ , and  $U$  as functions of  $k/\bar{M}_{Pl}$  when the SM resides on the TeV-brane. From bottom to top the curves correspond to  $m_1^{grav} = 200, 300, 400, 500, 750$ , and  $1000$  GeV.

# Present Constraints on RS Model

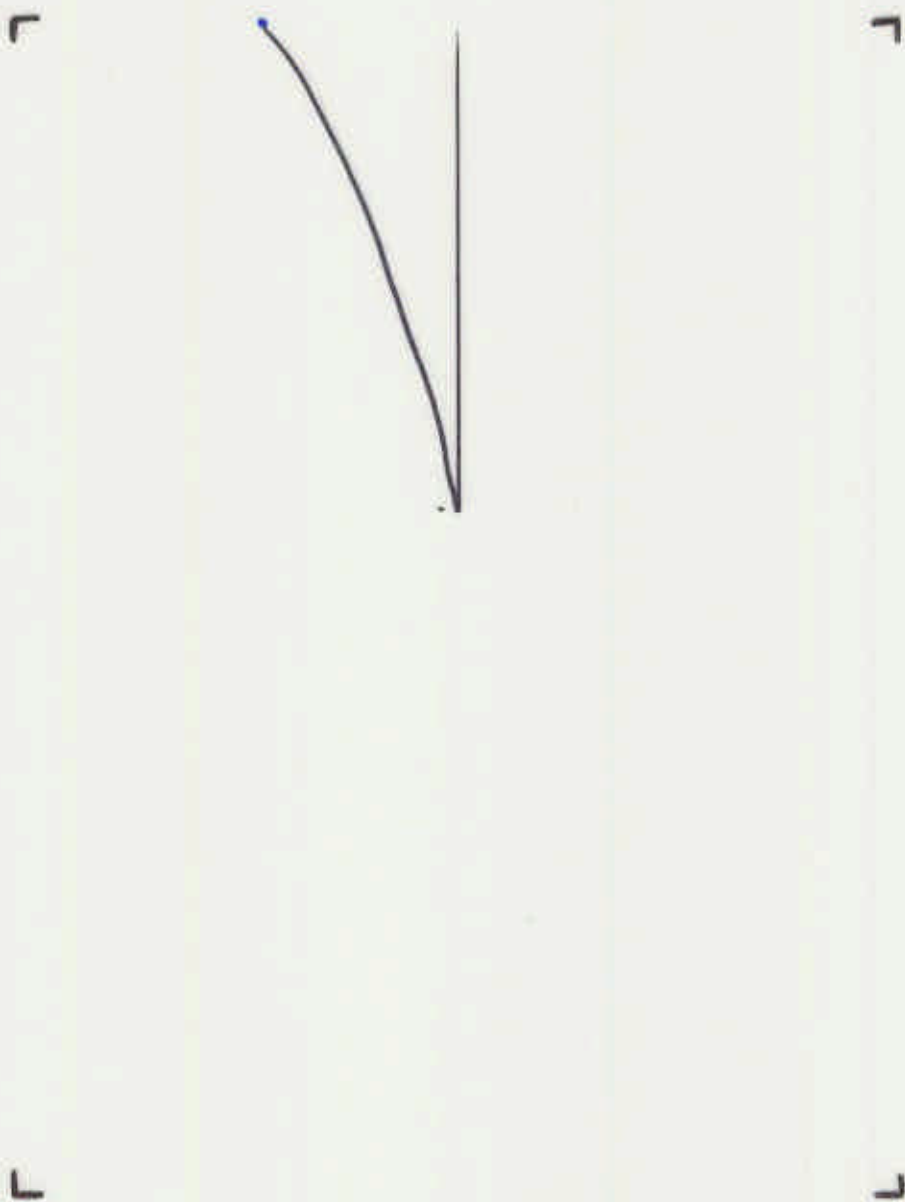


# Exp't + Theory Constraints





⇒ No-lose scenario for LHC!



## Peeling the SM Off the Wall

1) Gauge bosons in the bulk

(DHR  
Pomarol)

$$S_A = -\frac{1}{4} \int d^5x \sqrt{-G} G^{MK} G^{NL} F_{KL} F_{MN}$$

$$\text{KK expand: } A_M(x, \vartheta) = \sum_n A_M^{(n)}(x) \frac{\chi_A^{(n)}(\vartheta)}{\sqrt{r_c}}$$

$$\chi_A^{(n)}(\vartheta) = \frac{e^\sigma}{N_A} \left[ J_1\left(\frac{m_n}{k} e^\sigma\right) + \alpha_n Y_1\left(\frac{m_n}{k} e^\sigma\right) \right]$$

$$\mathcal{L}_{\text{FFA}} \sim g \bar{\psi} \gamma^\mu \psi \left[ A_\mu^{(0)}(x) + \sqrt{2\pi k r_c} \sum_{n=1}^{\infty} A_\mu^{(n)}(x) \right]$$

$$\text{For } \Lambda_\pi = e^{-k r_c \pi} \bar{m}_{p1} \sim \text{TeV}$$

$$g^{(n)} = \sqrt{2\pi k r_c} g^{(0)} \sim (8-9) g^{(0)} !$$

⇒ Tough constraints from Precision EW Data!

Data constrains  $m_i^A > 23 \text{ TeV}$

For  $\Lambda_\pi \sim 1 \text{ TeV}$  violates curvature constraint  
 $|R_5| = 20k^2 < m_5^2$

<OR>

For  $|R_5| < m_5^2$   $\Lambda_\pi > 100 \text{ TeV}$

## 2) Fermions in the bulk

$$S_f = \int d^5x \sqrt{G} \left[ e^\sigma \delta_n^m \left( \frac{i}{2} \bar{\psi} \gamma^n \partial_m \psi + \text{h.c.} \right) - \text{sgn}(\sigma) m \bar{\psi} \psi \right]$$

Note:  $\psi(-y) = \pm \gamma_5 \psi(y) \Rightarrow \bar{\psi} \psi$  is odd under  $Z_2$

$$\text{KK expand: } \psi_{L,R}(x, \vartheta) = \sum_n \psi_{L,R}^{(n)}(x) \frac{e^{2\sigma}}{\sqrt{r_c}} f_{L,R}^{(n)}(\vartheta)$$

integrate over  $\vartheta$  + impose orthonormality

$$\Rightarrow f_{L,R}^{(n)}(\vartheta) = \frac{e^{\sigma/2}}{N_n^{L,R}} \left[ J_{\nu/2 \mp \nu}(z_n^{L,R}) + \beta_n^{L,R} Y_{\nu/2 \mp \nu}(z_n^{L,R}) \right]$$

Choose  $f_R^{(n)}$   $Z_2$ -odd

$f_L^{(n)}$   $Z_2$ -even  $\Rightarrow f_L^{(0)} = \frac{e^{\nu\sigma}}{N_0^L} \equiv \text{SM fermions}$

Find  $\nu \gtrsim -0.9/-0.8$  to reproduce SM Yukawa's

★ Introduces an additional parameter

$$m_{\text{bulk}} = \nu k \quad ; \quad \nu \sim \mathcal{O}(1)$$

Find zero-modes couple more weakly than  
Wall fields

$\Rightarrow$  Serious reduction in collider sensitivity  
to RS phenomena!

## Coupling Coefficients

$f^{(0)} \bar{f}^{(0)} A^{(n)}$ :

$$C_{00n}^{f\bar{f}A} = \frac{g^{(n)}}{g^{SM}} = \sqrt{2\pi k r_c} \left[ \frac{1+2\nu}{1-e^{2\nu+1}} \right] \int_{\epsilon}^1 dz z^{2\nu+1} \frac{J_1(x_n^A z) + \alpha_n^A Y_1(x_n^A z)}{|J_1(x_n^A) + \alpha_n^A Y_1(x_n^A)|}, \quad (51)$$

$f^{(0)} \bar{f}^{(0)} G^{(n)}$ :

$$C_{00n}^{f\bar{f}G} = \frac{1}{\epsilon} \left[ \frac{1+2\nu}{1-e^{2\nu+1}} \right] \int_{\epsilon}^1 dz z^{2\nu+2} \frac{J_2(x_n^G z)}{|J_2(x_n^G)|}, \quad (52)$$

$A^{(0)} A^{(0)} G^{(n)}$ :

$$C_{00n}^{AAG} = \frac{1}{\epsilon} \frac{2(1 - J_0(x_n^G))}{\pi k r_c (x_n^G)^2 |J_2(x_n^G)|}, \quad (53)$$

$f^{(\ell)} \bar{f}^{(0)} A^{(n)}$ :

$$C_{\ell 0n}^{f\bar{f}A} = \sqrt{2\pi k r_c} \left| \frac{2(1+2\nu)}{1-e^{2\nu+1}} \right|^{1/2} \int_{\epsilon}^1 dz z^{\nu+3/2} \frac{J_f(x_\ell^f z)}{|J_f(x_\ell^f)|} \frac{J_1(x_n^A z) + \alpha_n Y_1(x_n^A z)}{|J_1(x_n^A) + \alpha_n Y_1(x_n^A)|}, \quad (54)$$

$f^{(\ell)} \bar{f}^{(0)} G^{(n)}$ :

$$C_{\ell 0n}^{f\bar{f}G} = \frac{1}{\epsilon} \left| \frac{2(1+2\nu)}{1-e^{2\nu+1}} \right|^{1/2} \int_{\epsilon}^1 dz z^{\nu+5/2} \frac{J_f(x_\ell^f z)}{|J_f(x_\ell^f)|} \frac{J_2(x_n^G z)}{|J_2(x_n^G)|}, \quad (55)$$

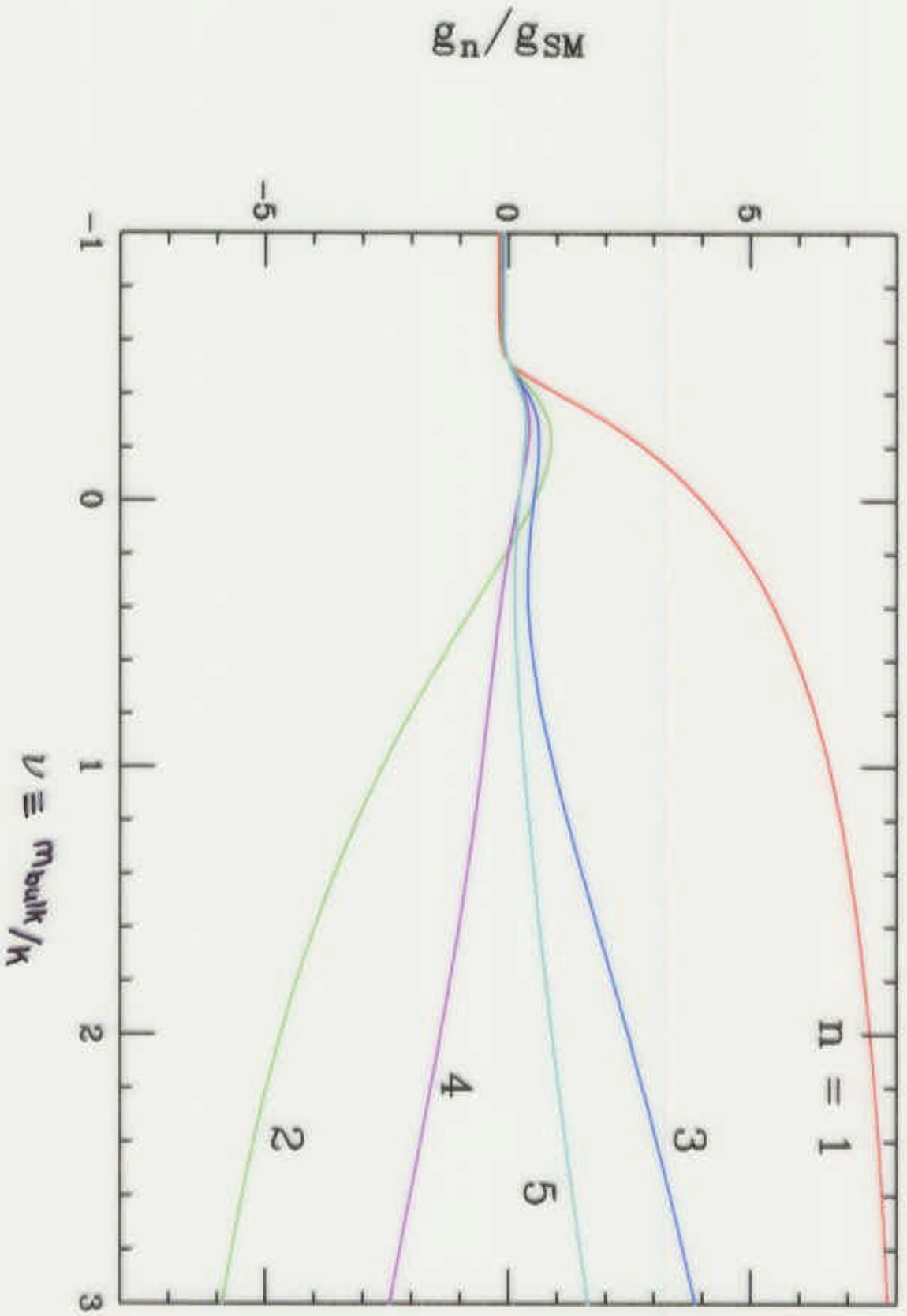
$A^{(\ell)} A^{(0)} G^{(n)}$ :

$$C_{\ell 0n}^{AAG} = \frac{2}{\epsilon \sqrt{2\pi k r_c}} \int_{\epsilon}^1 dz z^2 \frac{J_1(x_\ell^A z) + \alpha_\ell^A Y_1(x_\ell^A z)}{|J_1(x_\ell^A) + \alpha_\ell^A Y_1(x_\ell^A)|} \frac{J_2(x_n^G z)}{|J_2(x_n^G)|}, \quad (56)$$

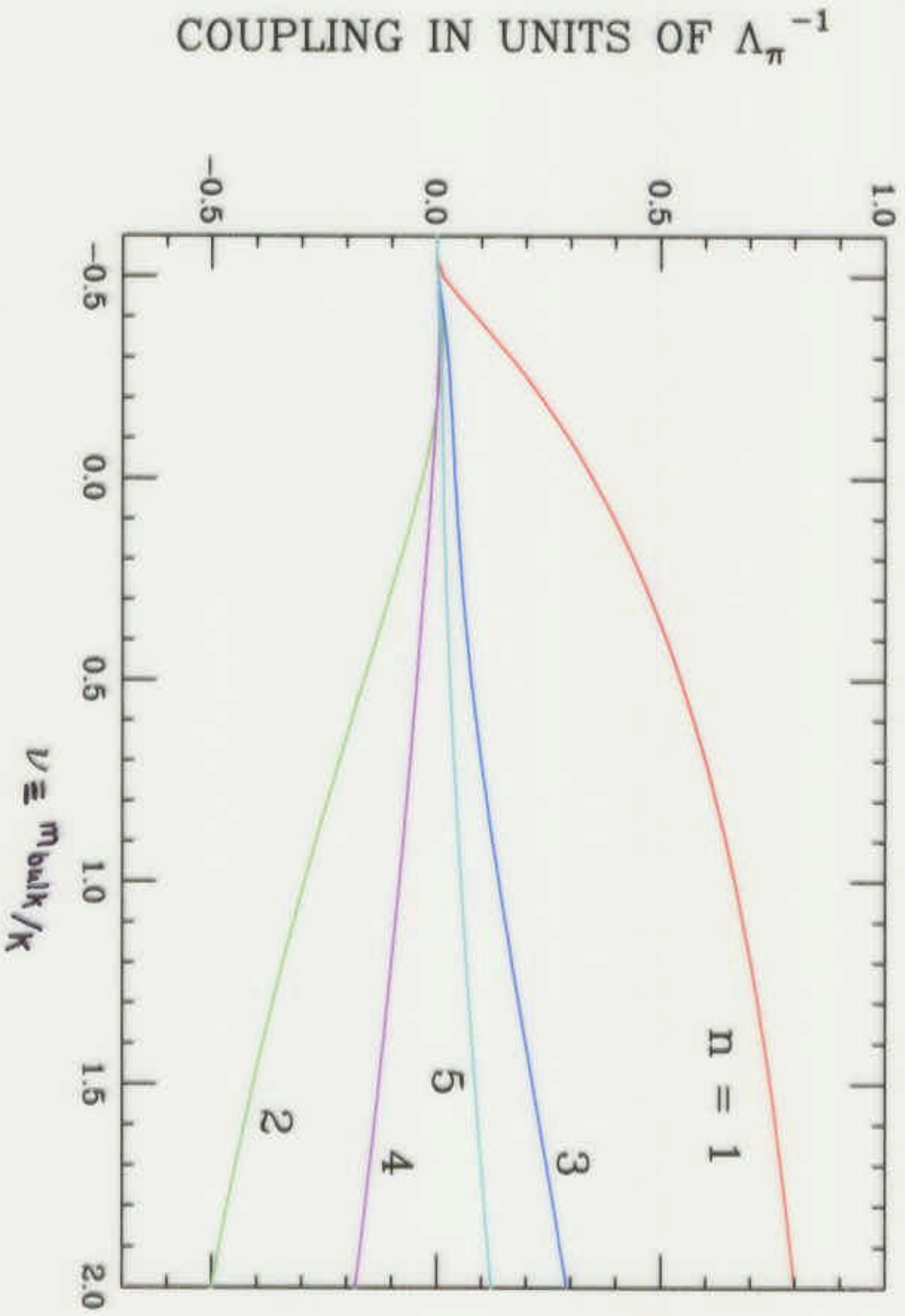
$f^{(\ell)} \bar{f}^{(0)} A^{(0)} G^{(n)}$ :

$$C_{\ell 0n}^{f\bar{f}AG} = \frac{1}{\epsilon} \left| \frac{2(1+2\nu)}{1-e^{2\nu+1}} \right|^{1/2} \int_{\epsilon}^1 dz z^{\nu+5/2} \frac{J_f(x_\ell^f z)}{|J_f(x_\ell^f)|} \frac{J_2(x_n^G z)}{|J_2(x_n^G)|}, \quad (57)$$

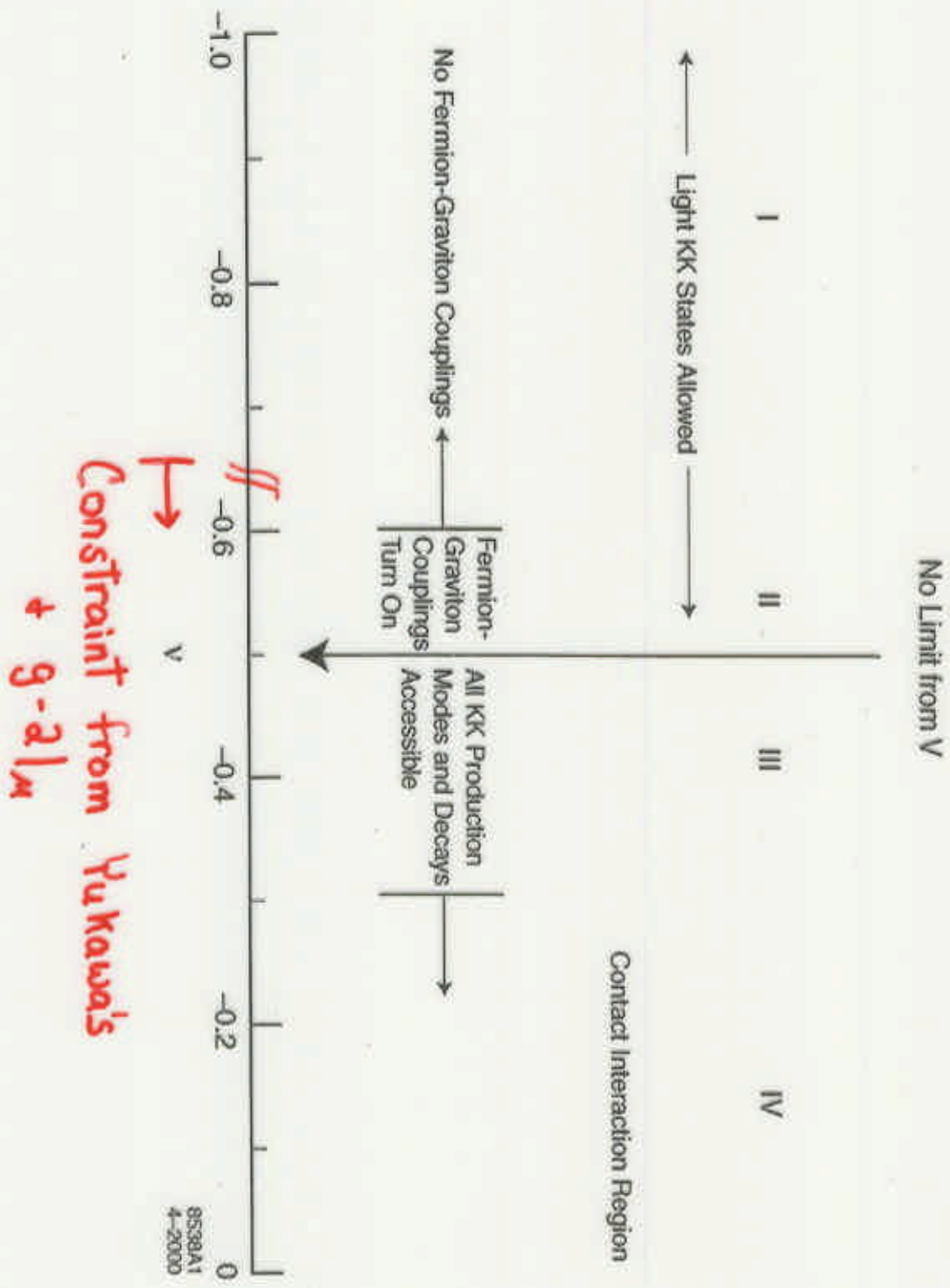
# KK Gauge Tower couplings to zero-mode fermions



# KK Graviton Tower Couplings to zero-mode fermions

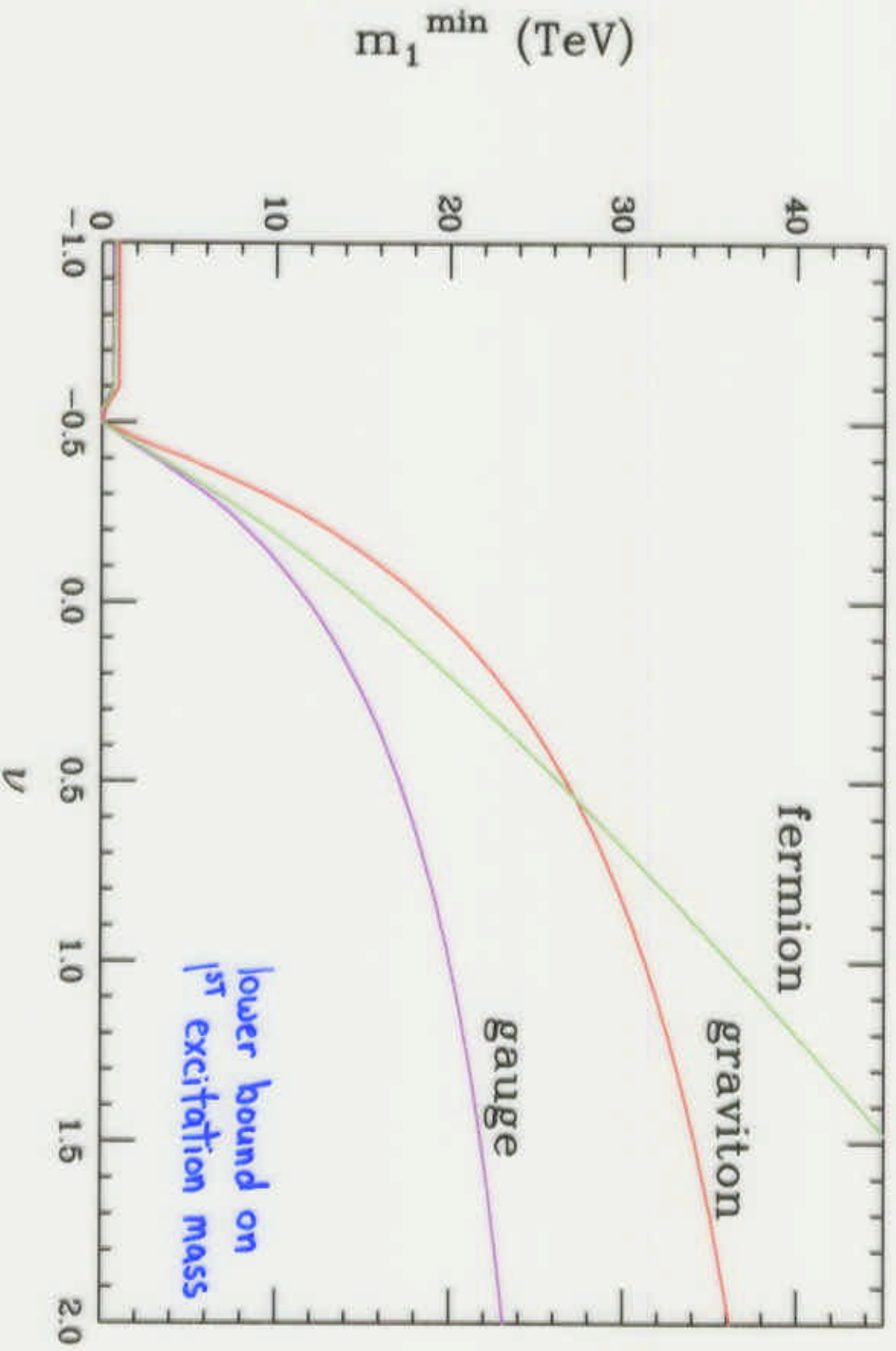


# Suggests Distinct Phenomenological Regions



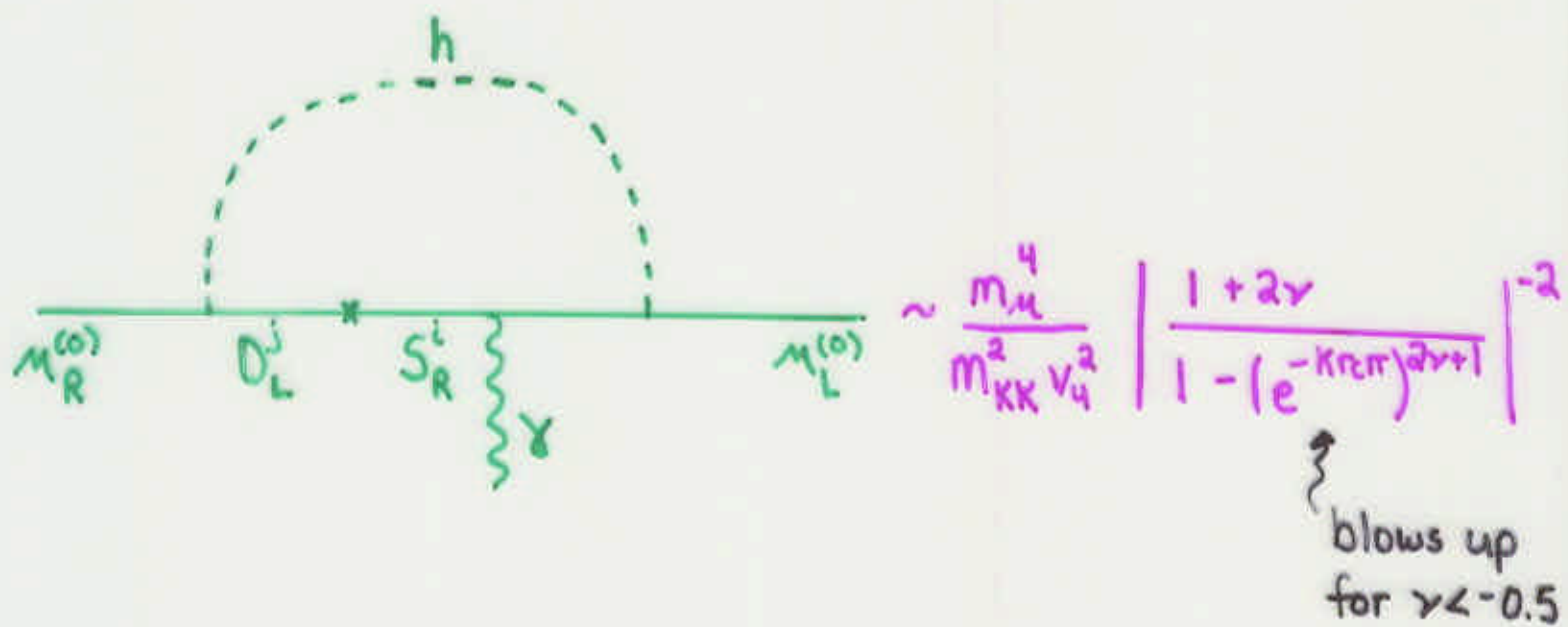
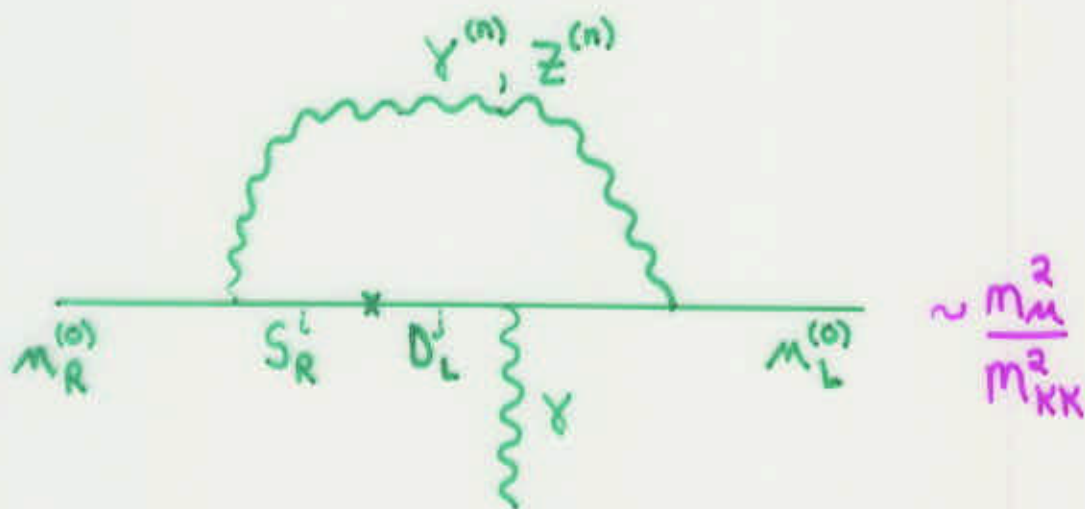
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## Constraints from Precision EW Data

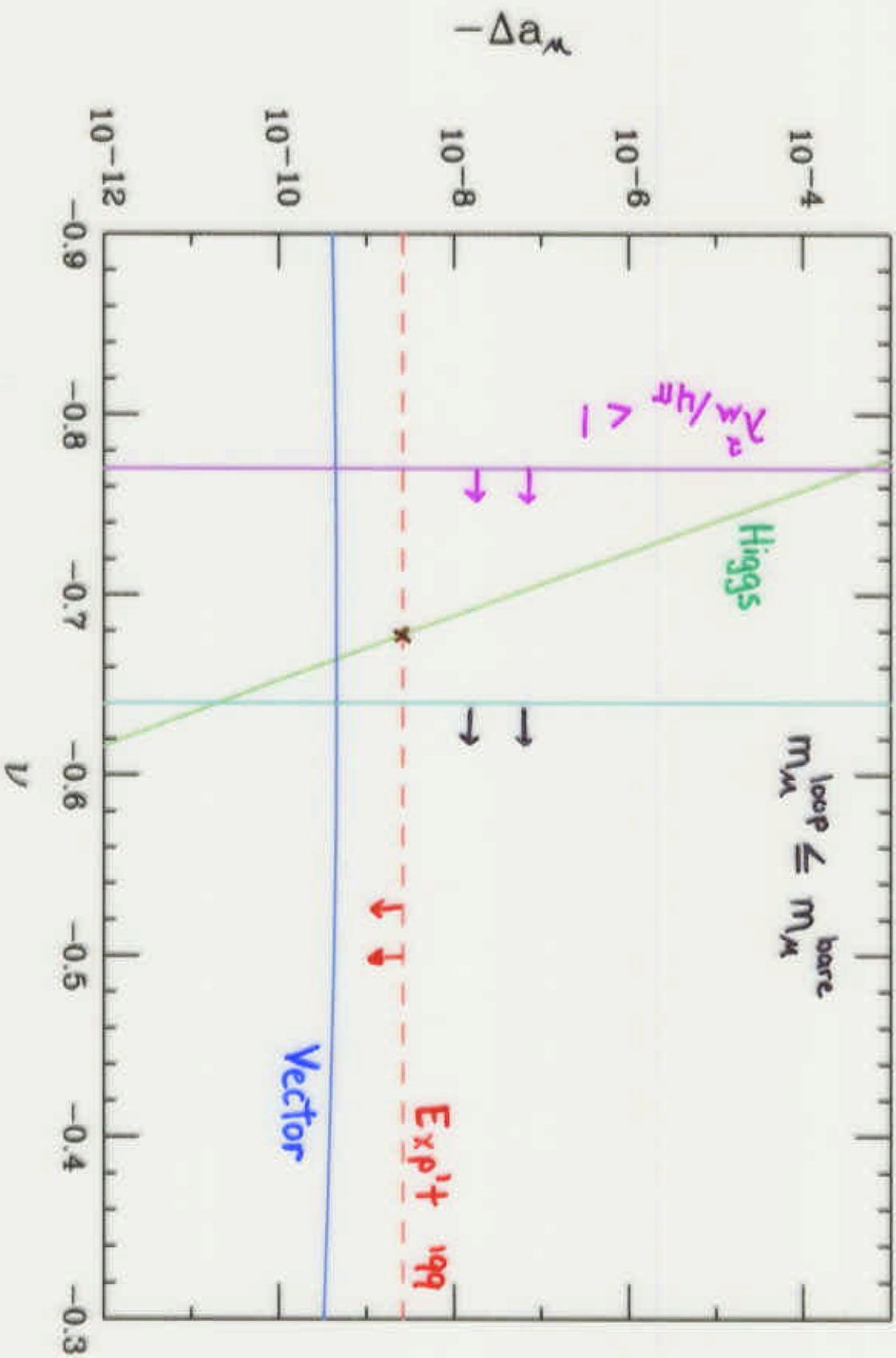




# $g=2|_m$ : Dominant Contributions



All other contributions further suppressed by  $m_M^n / m_{KK}^n$



## Conclusions

- Extra Dimensions provide novel approach to hierarchy problem!
- Two classes of theories - each with distinctive phenomenological tests at the weak scale

Large Extra Dimensions: predict

change in  $F_{\text{grav}}$  at  $d \lesssim 1 \text{ mm}$

change in  $f\bar{f}$  production +  $\cancel{E}_T$  dist's at colliders

Warped Extra Dimensions: predict

Graviton resonances at colliders

No-lose theorem @ LHC!