

EXTRA DIMENSIONS

AND

BENDING OF LIGHT

BRUCE MCKELLAR

GIRISH JOSHI

XIAO-GANG HE *

} UNIVERSITY
OF
MELBOURNE

* N. T. U.

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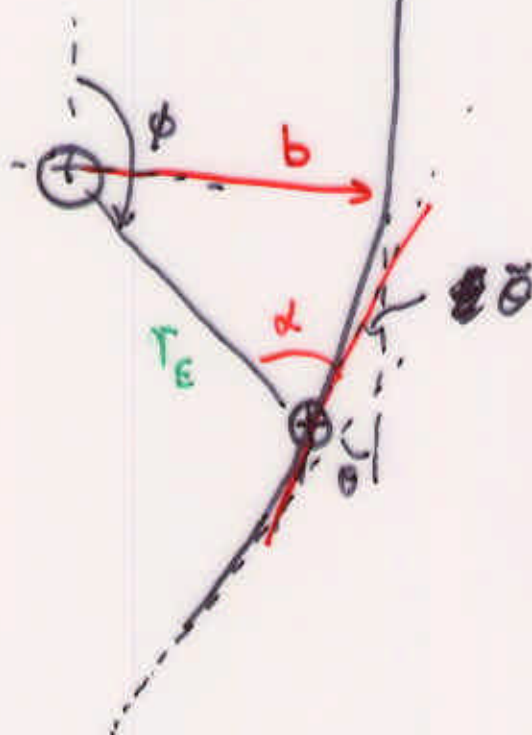
OSAKA

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THE CLASSIC CALCULATION

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Light bending in GR.



$$\alpha = \pi - \phi_E + \delta\alpha$$

observed angle
between star & Sun.

The geodesic is

$$\frac{b}{r} = \sin \phi + G \frac{(1+\gamma)M_\odot}{b} (1 - \cos \phi)$$

$$\phi(r=\infty) = \pi + \theta$$

$$\theta = \frac{2(1+\gamma)M_\odot G}{b} = \left(\frac{1+\gamma}{2}\right) \times 1.75''$$

$$\text{at } b = r_\odot$$

The deflection angle is

$$\tilde{\theta} = (1 + \gamma) \frac{M_{\odot} G}{r_E} \left(\frac{1 + \cos \alpha}{1 - \cos \alpha} \right)^{1/2}$$

at $\alpha = \frac{\pi}{2}$, $\tilde{\theta} = (1 + \gamma) \frac{M_{\odot} G}{r_E} \sim \frac{1 + \gamma}{2} 0''.0041$

THE MOST ACCURATE CURRENT DATA is

FOR $40^{\circ} \leq \alpha \leq 140^{\circ}$

[HIPPARCOS SATELLITE]

RESULTS

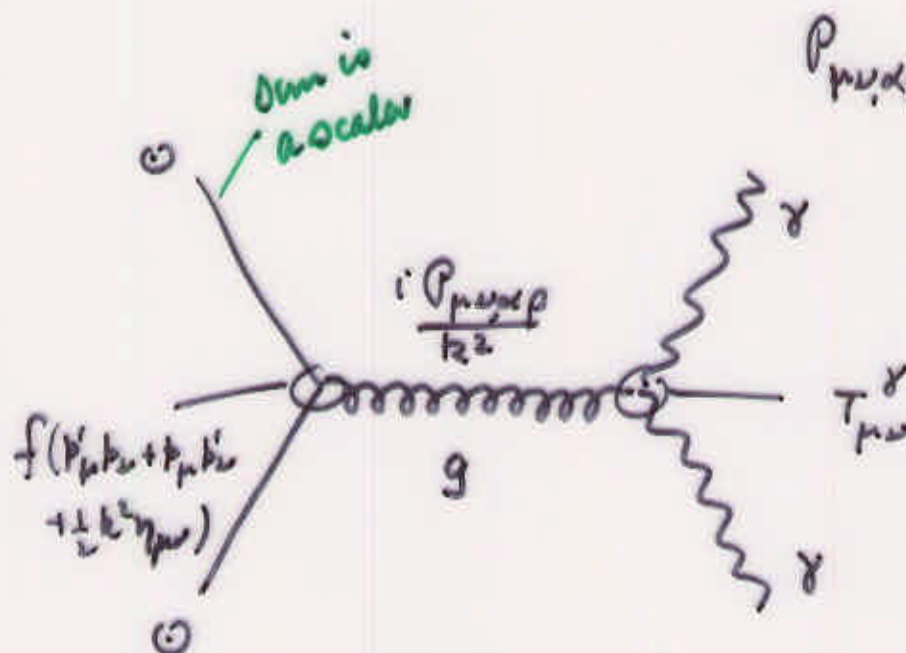
$$\gamma = 0.997 \pm 0.003$$

GR $\gamma = 1$ ✓

ALTERNATIVE CALCULATION!

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GRAVITON EXCHANGE = LIGHT BENDING.



$$P_{\mu\nu,\alpha\beta} = \eta_{\mu\alpha}\eta_{\nu\beta} + \eta_{\nu\alpha}\eta_{\mu\beta} - \eta_{\mu\nu}\eta_{\alpha\beta}$$

$$T_{\mu\nu}^{\gamma} = F_{\mu}^{\alpha} F_{\nu\alpha} + \frac{1}{4} \eta_{\mu\nu} F^{\alpha\beta} F_{\alpha\beta}$$

$$= \epsilon^{\alpha\beta} T_{\mu\nu\alpha\beta}^{\gamma} \epsilon^{\gamma}$$

$$T_{\mu\nu,\alpha\beta}^{\gamma} = \frac{1}{2} [q'_{\alpha} (q_{\mu} \eta_{\nu\beta} + q_{\nu} \eta_{\alpha\mu}) + q_{\beta} (q'_{\mu} \eta_{\nu\alpha} + q'_{\nu} \eta_{\mu\alpha})$$

$$- \eta_{\alpha\beta} (q'_{\mu} q_{\nu} + q_{\mu} q'_{\nu}) + \eta_{\mu\nu} (q'_{\alpha} q_{\beta} - q_{\alpha} q'_{\beta})$$

$$- q'_{\alpha} q_{\beta} (\eta_{\mu\alpha} \eta_{\nu\beta} + \eta_{\nu\alpha} \eta_{\mu\beta})]$$

$$S_{fi} = (-i)^2 4 f^2 p'_s p_r i \frac{\rho^{\delta\sigma\mu\nu}}{k^2}$$

$$[g'_\alpha g_\mu \eta_{\rho\nu} + g_\rho g'_\mu \eta_{\alpha\nu} - \eta_{\alpha\rho} g'_\mu g_\nu - \frac{1}{4} \eta_{\alpha\rho} g_\nu g'_\alpha] \epsilon^{\rho\sigma}(\mathbf{q}') \epsilon^\nu(\mathbf{q})$$

$$\rightarrow \frac{1}{2} \left| \frac{f^2 M_\odot^2}{\sin^2 \frac{\theta}{2}} \right|^2 \sum_{\text{pol}} |\epsilon' \cdot \epsilon|^2 = \left| \frac{f^2 M_\odot^2}{\sin^2 \frac{\theta}{2}} \right|^2$$

$$\frac{d\sigma}{d\Omega} = \left| \frac{T_{fi}}{8\pi M_\odot} \right|^2 = \frac{16 G^2 M_\odot^2}{\theta^4}$$

$$b db = - \frac{d\sigma}{d\Omega} \cdot \sin \theta d\theta$$

$$2 \int_{b_{\min}}^{b_{\max}} b db = - \int_{\theta_i}^{\bar{\theta}} \frac{d\sigma}{d\Omega} \theta d\theta = - 16 G^2 M_\odot^2 \int_{\theta_i}^{\bar{\theta}} \frac{d\theta}{\theta^3}$$

$$2 R_\odot^2 = 16 G^2 M_\odot^2 \frac{1}{2 \bar{\theta}^2}$$

$$\bar{\theta} = \frac{4 G M_\odot}{R_\odot}$$

At $b = R_\odot$

as usual

WHICH TECHNIQUE DO WE USE FOR
GENERALISING TO EXTRA DIMENSIONS?

GRAVITON EXCHANGE HAS BEEN USED
TO INTRODUCE 1 LOOP CORRECTIONS
TO EINSTEIN'S RESULT

FOLLOW THIS WAY TO EXTRA
DIMENSIONS — EXCHANGE
KALUZA KLEIN TOWERS AS
WELL AS GRAVITONS.

GENERALISATION

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With Extra Dimensions

Now allow the ~~metric~~
to represent

- graviton
- KK towers of $s_{\mu\nu}$ 0, 1, 2

But only $s_{\mu\nu}$ 2 couples to both scalar
matter field and photon.

Additional KK propagator term

$$\sum_l \frac{B_{\mu\nu,\alpha\rho}^{KK}(k)}{k^2 - m_l^2}$$

$$B_{\mu\nu,\alpha\rho}^{KK}(q) = \left(\eta_{\mu\nu} - \frac{q_\mu q_\nu}{m_l^2} \right) \left(\eta_{\alpha\rho} - \frac{q_\alpha q_\rho}{m_l^2} \right) \\ + (\mu \leftrightarrow \nu) - \frac{2}{3} \left(\eta_{\mu\alpha} - \frac{q_\mu q_\alpha}{m_l^2} \right) \left(\eta_{\nu\rho} - \frac{q_\nu q_\rho}{m_l^2} \right)$$

Only $\eta_{\mu\alpha} \eta_{\nu\beta} + \eta_{\mu\beta} \eta_{\nu\alpha}$ terms survive
 [Gauge Invariance]

Just replace $\frac{1}{k^2} \rightarrow \frac{1}{k^2} + \sum_{\ell} \frac{1}{k^2 - m_{\ell}^2}$

The density of states are very large

$$\Delta = \sum_{\ell} \frac{1}{q^2 - m_{\ell}^2} = -\frac{2}{M_S^2 G_N} \left(\frac{|q|^2}{M_S^2} \right)^{\frac{n}{2}-1}$$

$$I_n(M_S, \sqrt{|q|^2})$$

$$I_n = \int_{M_{\min}/\sqrt{|q|^2}}^{M_S/\sqrt{|q|^2}} \frac{y^{n-1}}{\sqrt{1+y^2}} dy$$

$$M_{\min} = m_i \quad \text{KK norm} \quad \frac{2\pi}{R} \sim 10^{-2} \text{eV}$$

$$R \sim 10^{-2} \text{m.}$$

$$\Delta = \frac{1}{M_s^4 G_N} \ln \frac{M_s^2}{|q|^2 + M_{in}^2} \quad n=2$$

$$\Delta = \frac{1}{M_s^4 G_N} \frac{2}{n-2}$$

used $G_N = (4\pi)^{n/2} \Gamma(n/2) R^{-n} M_s^{-(n+2)}$

REPEAT CROSS SECTION CALCULATION:

$$\frac{d\sigma}{d\Omega} = 16 G_N^2 m^2 \left(\frac{1}{\theta^2} + \omega^2 \Delta \right)^2$$

AND BENDING ANGLE CALCULATION:

$$\theta = \frac{4 G_N M_\odot}{b} \left(1 - 2\omega^2 \Delta \cdot \left(\frac{4 G_N M_\odot}{b} \right)^2 \ln \frac{4 G_N M_\odot}{b} \right)$$

GR + Correction due to Extra Dimensions

Correction

- Strongly ω dependent - also true of 1 loop corrections
- $\propto \frac{1}{b^2}$

Only significant for $b \sim R_D$, where $\gamma = 1 + \Delta\gamma$

$$\Delta\gamma = -0.50 \left(\frac{\omega^2}{eV^2} \right) \left(\frac{1 \text{ TeV}}{M_s} \right)^4 \delta$$

$$\delta = \frac{2}{n-2} \quad n > 2$$

$$\equiv \ln \left(\frac{M_s^2}{\omega^2 b^2 + M_{pl}^2} \right) \quad n = 2$$

General result away from grazing incidence

$$\delta\alpha = (1 + \gamma) \frac{G_N M_\odot}{r_E} \frac{1}{\sin\alpha}$$

$$\gamma = 1 - 4w^2 \Delta \left(\frac{4G_N M_\odot}{r_E \sin\alpha} \right)^2 \approx \left(\frac{4G_N M_\odot}{r_E \sin\alpha} \right)$$

The Hipparcos results are $\sin\alpha \sim 1$

$$\delta\alpha \ll 1 \text{ — no constraint}$$

Go back to eclipse data $r_E \sin\alpha = r_\odot$

$$\gamma = 0.95 \pm 0.11$$

$$M_S \geq 1.4 \left(\frac{2}{n-2} \right)^{1/4} \text{ TeV} \quad n > 2$$

$$\sim 4.2 \text{ TeV} \quad n \sim 2.$$

OPEN QUESTIONS

1. EXPERIMENTAL

CAN WE GET 10% ACCURACY
AT $\omega \approx 1 \text{ MeV}$

THIS WOULD PERMIT 10^3 TeV LIMITS
ON M_s

2. THEORETICAL

CALCULATING THE KK CONTRIBUTION
TO THE POTENTIAL, AND INSERTING
THAT INTO THE LINEARISED
EINSTEIN EQUATIONS GIVES A
NEGLECTIBLE CORRECTION — SUPPRESSED
BY $e^{-M_s R_3}$

A GOOD THEORETICAL ARGUMENT
PREFERRING ONE CALCULATION OVER
THE OTHER IS NEEDED.