



ICHEP 2000



F Ronga
INFN Frascati

Atmospheric neutrino results from:



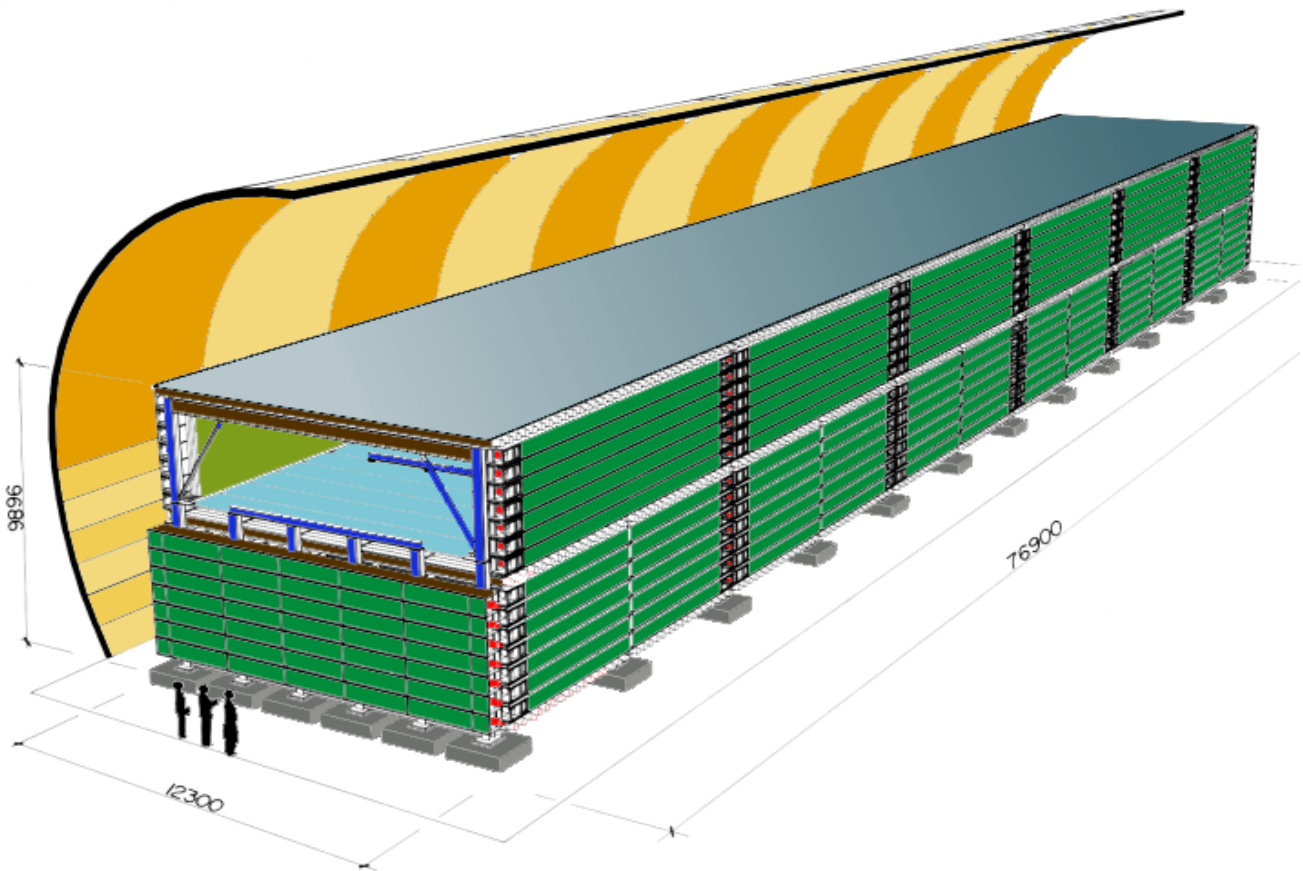
Monopole , **A**strophysics , and **C**osmic **R**ay **O**bservatory

*Bari, Bologna, Boston, Caltech, Drexel, Indiana, Frascati,
Gran Sasso , L'Aquila, Lecce, Michigan, Napoli, Pisa, Roma I,
Texas, Torino*

Summary:

- Upward- Going Through-Going Muons
- Sterile Vs Tau Neutrino with matter effects
- Low energy Events
- Conclusions

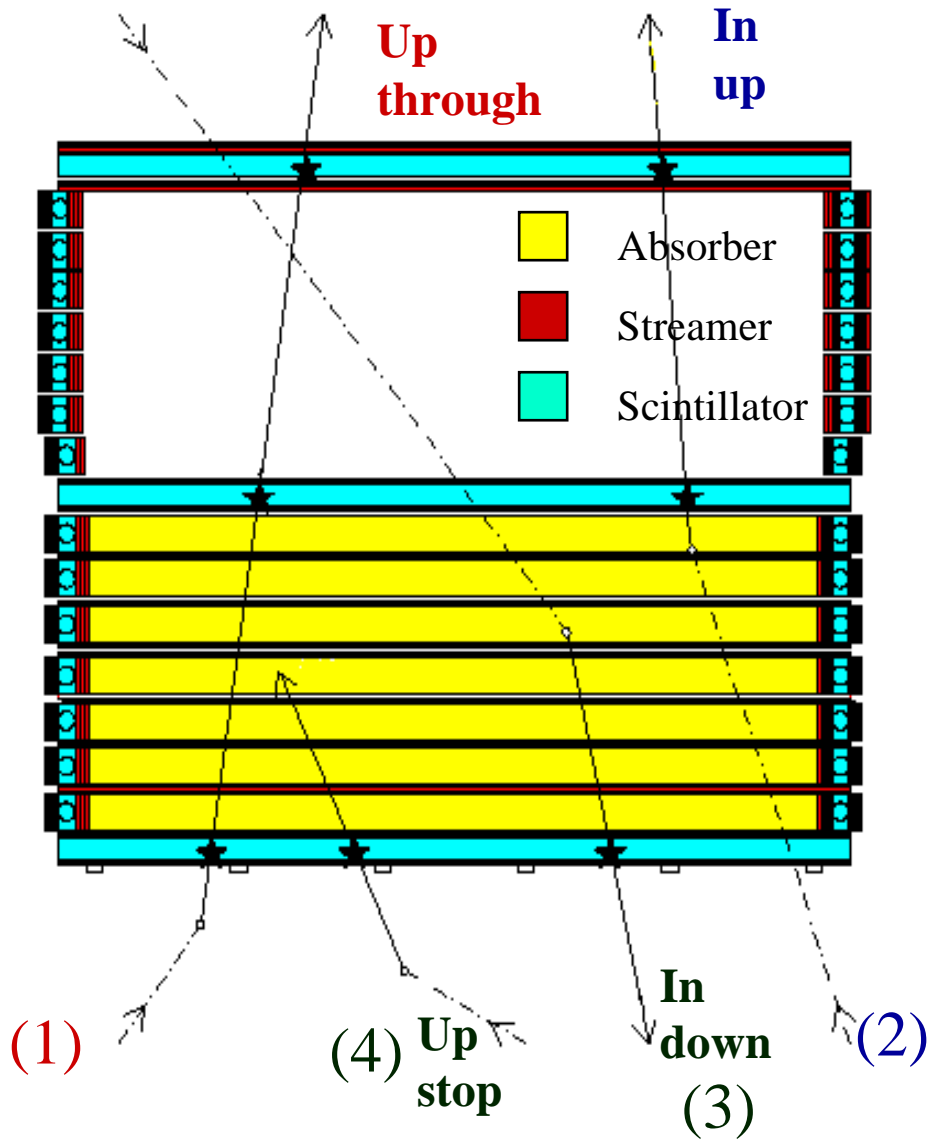
Main features of Macro as ν detector



- Large acceptance ($\sim 10000 \text{ m}^2\text{sr}$ for an isotropic flux)
- Low downgoing μ rate ($\sim 10^{-6}$ of the surface rate)
- ~ 600 tons of liquid scintillator to measure T.O.F. (time resolution $\sim 500\text{psec}$)
- $\sim 20000 \text{ m}^2$ of streamer tubes (3cm cells) for tracking (angular resolution $< 1^\circ$)

More details in Nucl. Inst. and Meth. A324 (1993) 337.

Neutrino event topologies in MACRO



Detector mass ~ 5.3 kton

(1) Up throughgoing μ (ToF) ~ 140 Ev/y

$E_{\text{median}} \approx 50$ GeV

(2) Internal Upgoing μ (ToF) ~ 25 /y $E_{\text{median}} \approx 3.5$ GeV

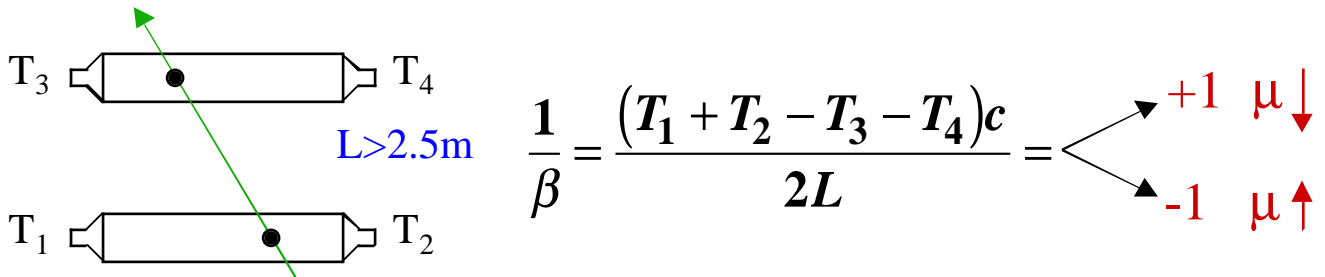
(3) Internal Downgoing μ (no ToF) ~ 22 y

$E_{\text{median}} \approx 4.2$ GeV

(4) Upgoing Stopping μ (no ToF) ~ 22 /y

Upward Going Muons (Through) Data Set

data selection based on the time-of flight measured with the scintillators



- 1) streamer track in agreement in the scintillator counters with the position from the times at the ends
- 2) $-1.25 < 1/\beta < -0.75$
- 3) 2m ($\approx 200\text{gr/cm}^2$ of absorber crossed to reduce at 1% level the backward due to π 's produced by downward-going muons (Astrop Phys 9 (1998) 105)

- Data selection only by software. (No scanning)

- 768 neutrino induced through-going upward muons during **Mar 89-Nov 91 (1/6 lower detector, 1.38 y)**

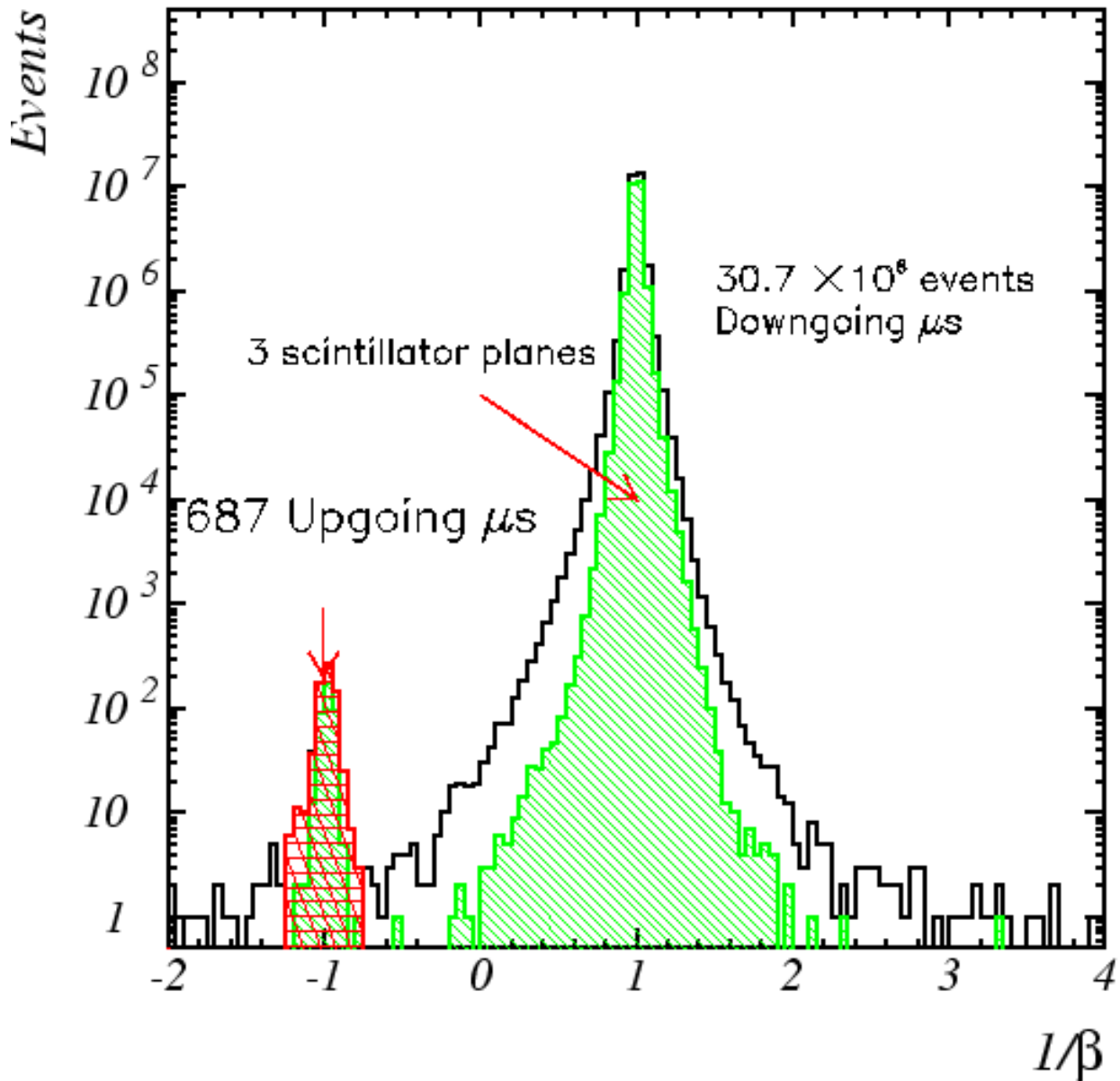
- Dec 92-Jun 93 (lower detector, 0.41 y)**

Phys. Lett. B357 (1995) 481

- 29 Apr 94 –5 Mar 2000 (complete detector) 4.8y**

(Phys.Lett. B434 (1998) 451 until Nov 98)

Upward Going Muons $1/\beta$ distribution

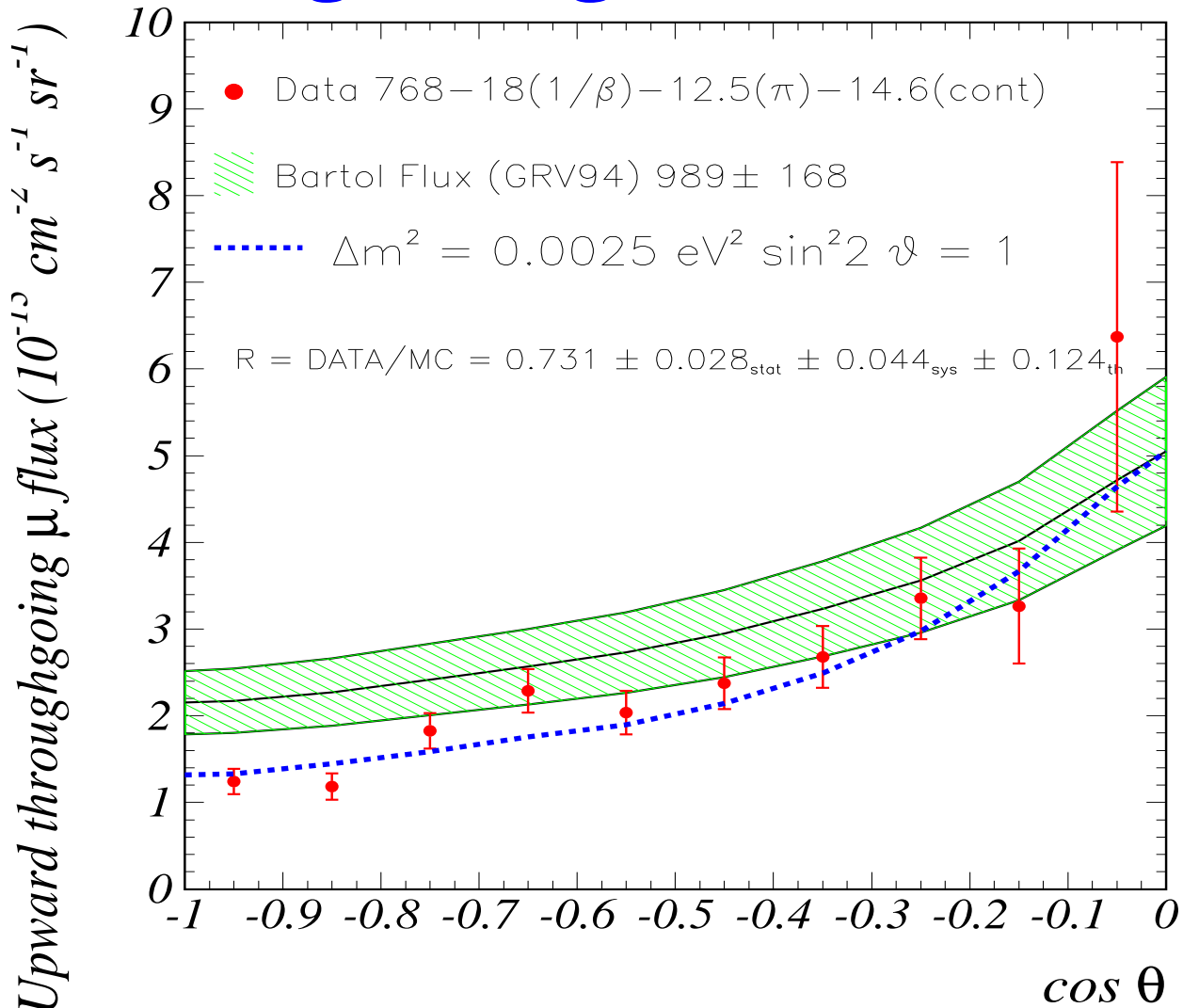


Upward Going Muons (Through) Results

Total number of events:	768
background (wrong b)	18
background (pion from muon)	12.5
Internal neutrino interactions	14.6
Total	723
Prediction	989\pm17%
Bartol neutrino Flux \pm 14%	
GRV94 cross section 9%	
Lohmann muon energy loss 5%	
R=data/prediction=	0.73

$\pm 0.028(\text{stat}) \pm 0.044(\text{systemat.}) \pm 0.12(\text{theoretical})$

Upward Going Muons (Through) Angular Distribution



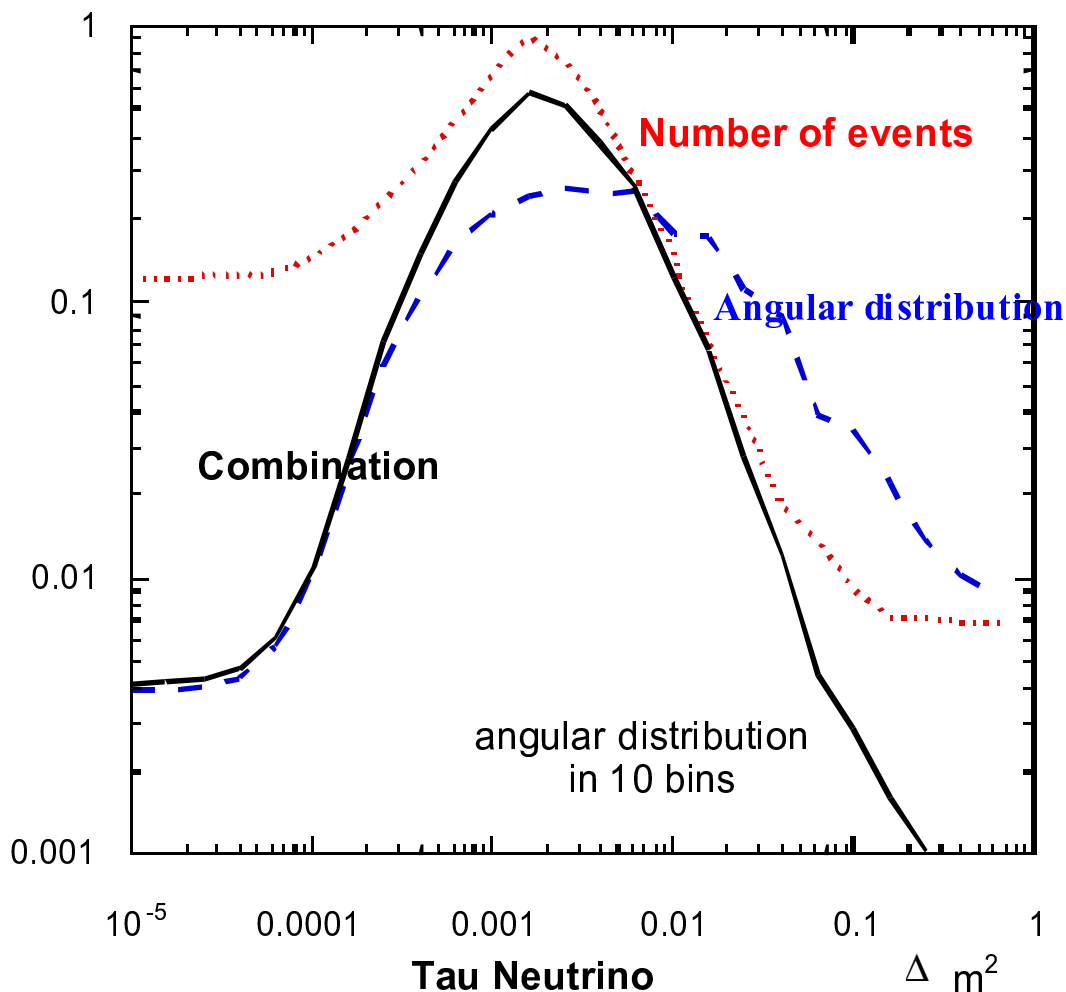
χ^2 test on the angular distribution (10 bins)
with prediction normalized to data :

- $\chi^2 = 11.2/9$ d.o.f. for $\nu_\mu \Rightarrow \nu_\tau$ with maximum mixing and $\text{dm}^2 \sim 0.0025 \text{ eV}^2$ $P = 26 \%$
- $\chi^2 = 24.3/9$ d.o.f. for no - oscillations $P = 0.9 \%$

(since $\langle E_\mu \text{ thresh} \rangle \approx 1 \text{ GeV}$ MACRO should be compared with the SuperKamiokande Through-going + Stopping muons)

Probabilities for maximum mixing and $\nu_{\mu} \rightarrow \nu_{\tau}$ oscillations

Peak probability from the angular distribution: 26%
from the combination: 57%



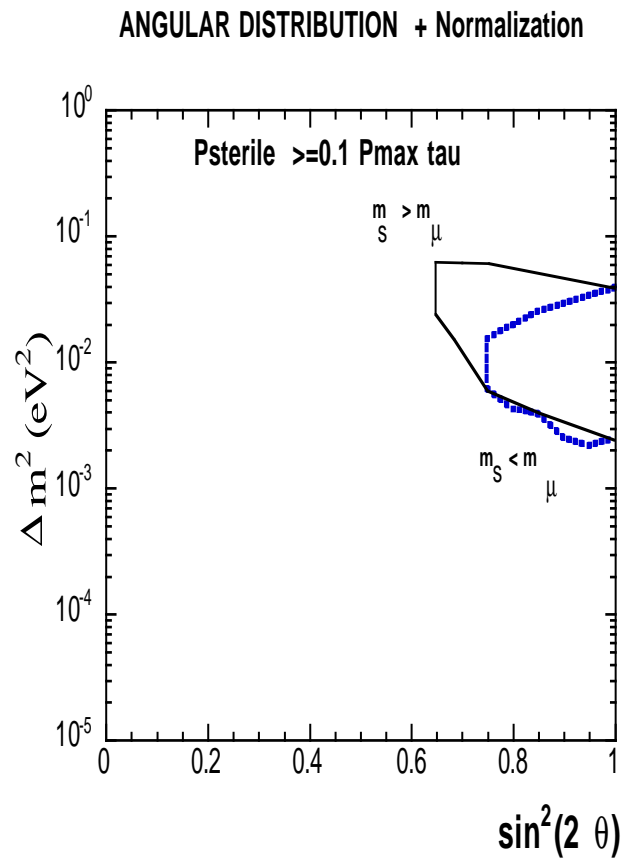
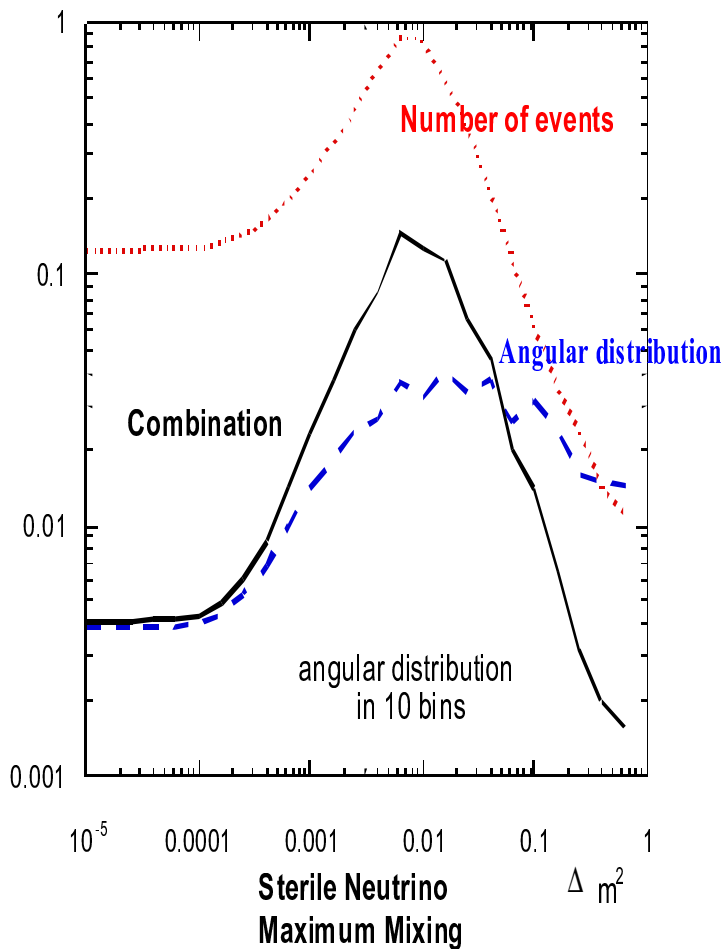
- The peak probability from the angular distribution alone is in the same region of the peak probability from the total number of events
- Probability for no-oscillation: $\sim 0.4\%$

MACRO UPMU

Probability for sterile neutrinos oscillations

- sterile neutrino \implies matter effects \implies reduction of the angular distortion

Peak probability from the angular distribution: 4.1%
 from the combination: 14.5%



Peak probabilities **lower than that for tau neutrinos:**

- from the angular distribution: 4.1 %
- from combination: 14.5 %

MACRO UPMU : matter effect with the ratio vertical/horizontal

- This ratio (*Lipari -Lusignoli (Ph Rev D 57 1998)*) can be statistically more powerful than a chi-square test for two reasons:

- 1) the ratio is sensitive to the sign of the deviation

- 2) there is gain in statistical significance grouping data in two bins

- As disadvantage you could lost **some data structure** in the angular distribution

- Ratio or chi-square in 10 bins ? Several authors prefer chi-square.

Chi-square in 10 (or more) bins **no strong discrimination between tau and sterile neutrino oscillations** (for SK also)

(Foot hep-ph/0007065 , Fornengo et al hep-ph/0002147)

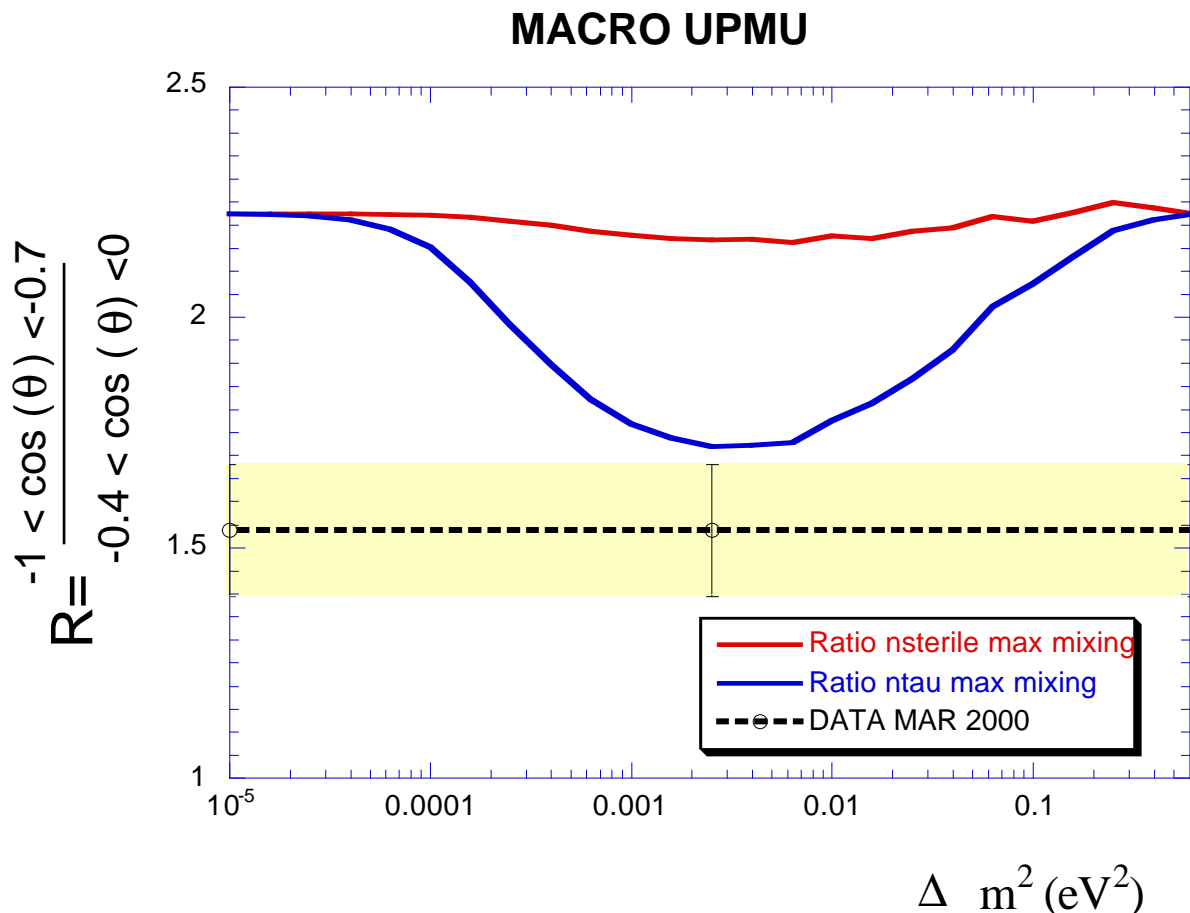
- **Recently : optimization of the ratio**

Result (for MACRO) for the best bin combination

$$R = \frac{N(\cos(\theta) < -0.7)}{N(\cos(\theta) > -0.4)}$$

*Obtained from a **Montecarlo simulation** to minimize the probability for sterile neutrino assuming tau neutrino oscillations with $dm^2=0.0025 eV^2$*

Ratio vertical/horizontal Montecarlo Optimized



P best Tau/ P best Sterile =
 70 (5% systematic in each bin)
 413 (no systematic error)

- **The plot is for Maximum mixing. But similar results for mixing < 1**
- **Sterile neutrino disfavored respect to tau at >98%** (5% systematic in each bin)

Errors on the ratio vertical/horizontal

- **neutrino flux (Lipari neutrino 2000) :**

the ratio depend from the K / π ratio and decay

$$L_{\text{decay } K} \approx 0.75 \text{ (EGeV/100) Km}$$

$$L_{\text{decay } \pi} \approx 5.6 \text{ (EGeV /100) Km}$$

two sources of uncertainty :

a) Spectral Index in the energy spectrum of the primary $E^{-\alpha}$

b) The calculation of the K/π fraction in the cosmic ray cascade
(largest contribution)

$$\frac{\delta R}{R} \approx 3\%$$

- **uncertainty in the neutrino cross section as function of the energy:**

from the comparison of different cross sections (MACRO)

$$\frac{\delta R}{R} \approx 2\%$$

- **uncertainty in the detector acceptance :**

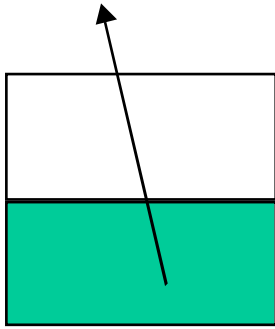
from a study of the down-going muons in MACRO we obtain 5% in each bin (used in the likelihood ratio calculation) ==> (if not correlated)

$$\frac{\delta R}{R} \approx 7.1\%$$

- **uncertainty in the background in the bin near the horizontal:**

Foot hep-ph/0007065

negligible in MACRO : small number of events and conservative cut (exclusion of 50% solid angle $\cos(\theta) > -0.1$)

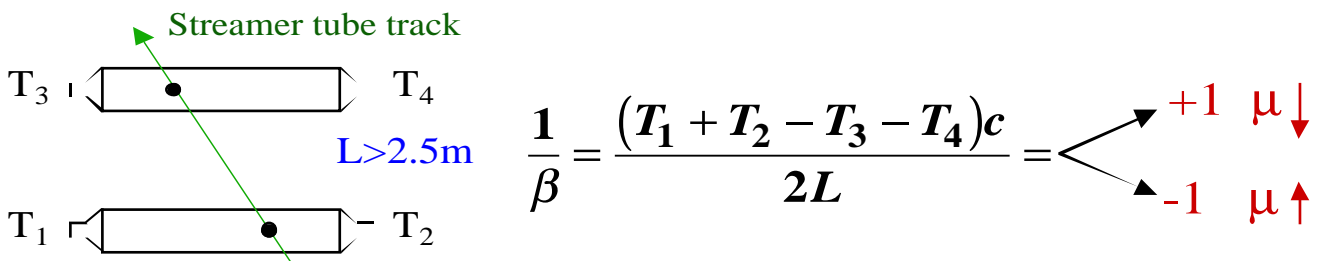


Internal Upgoing partially contained events (IU)

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Identified by:

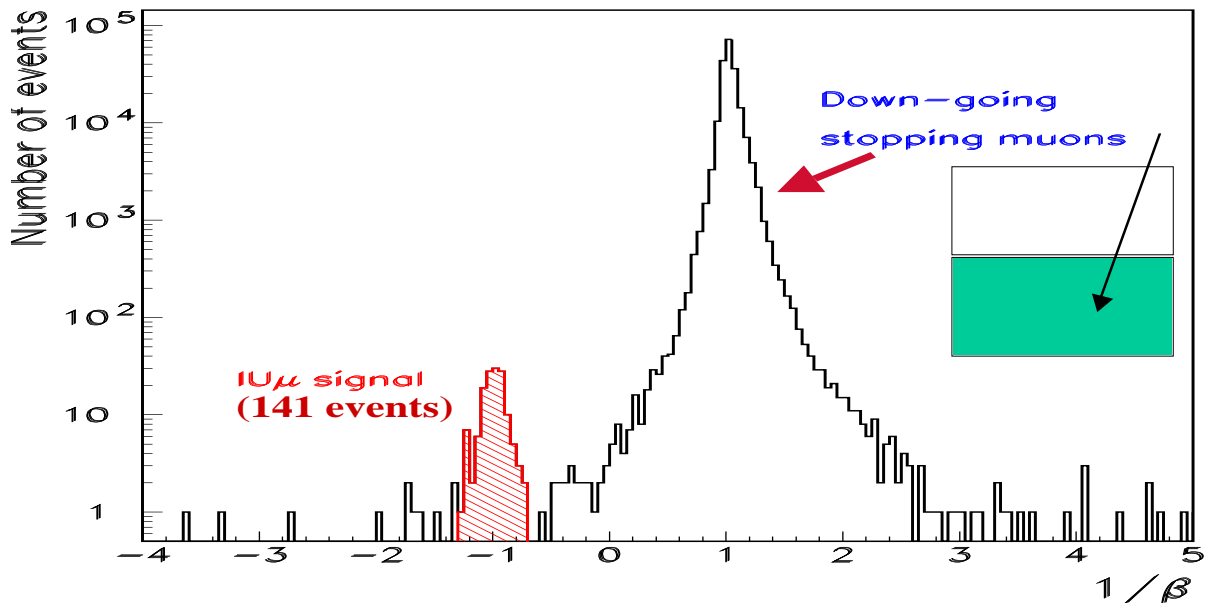
time-of-flight between central / upper SC layers



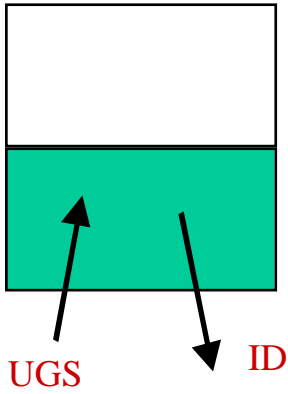
and **topological criteria for**

vertex containment inside lower detector

to remove upward-throughgoing muons (~1 % after this cut)



DATA - Bck($1/\beta$) = 135 events



Internal Downgoing (ID) and Upward Going Stopping (UGS) partially contained events

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Selection criteria:

- No T.o.F. measurement
- topological constraints for the track (bottom SC layer + track inside fiducial volume)
- visual scan procedure (on real and simulated events) for final selection
- $> 100 \text{ g cm}^{-2}$ of material crossed in the detector to reduce background of upward π from μ

From MC simulation:

- $E_\nu \sim 4 \text{ GeV}$
- Mixture of $\sim 50\%$ UGS μ + $\sim 50\%$ ID μ
- $\sim 87\%$ from ν_μ -C.C. interactions

DATA: 5.1 live-y

- From Apr. 1994 up to Mar. 2000

DATA - Bck = 229 events

MONTE CARLO PREDICTIONS

- GEANT based program for the simulation of detector response
- **Simulated events processed through the same analysis chain as the data**

$$\Phi_{\nu_{\mu} \rightarrow \mu} = \Phi_{\nu} \otimes \sigma_{\nu} \otimes \varepsilon(E_{\mu}, \theta_{zenith})$$

- Φ_{ν} : **Bartol ν flux** with geomagnetic cutoffs (error ~ 20%)
- $\sigma_{\nu} = \text{Q.E.} + 1\pi$ (Lipari et al., PRL74 (1995) 4384) + DIS (GRV-LO-94 PDF) (error ~ 15%)
- $\varepsilon(E_{\mu}, \theta_{zenith})$: detector response and acceptance (systematic error ~ 10 %)



Internal Up

MC: $247 \pm 25_{\text{sys}} \pm 62_{\text{the}}$
DATA: $135 \pm 12_{\text{stat}}$

Idown + UGStop

MC: $329 \pm 33_{\text{sys}} \pm 82_{\text{the}}$
DATA: $229 \pm 15_{\text{stat}}$

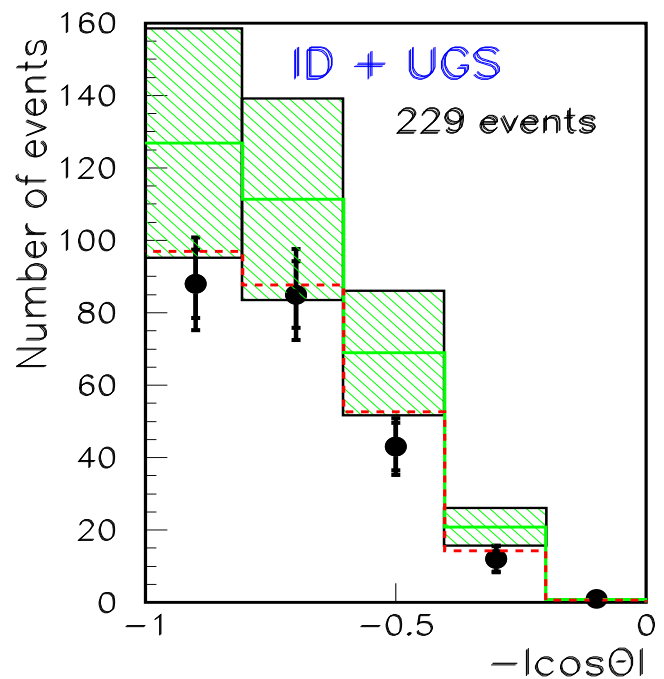
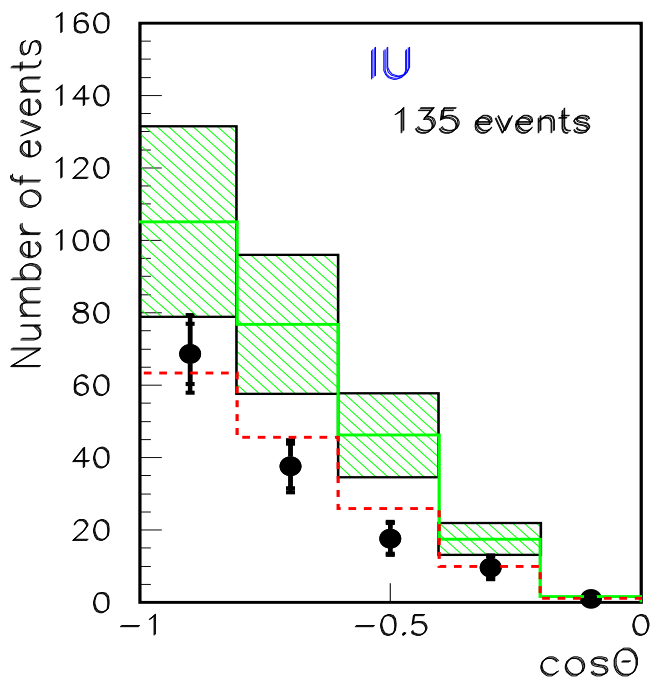
- Using the NEUGEN code (MINOS experiment):
→ 6% difference in the expectations

ANGULAR DISTRIBUTIONS

LOW ENERGY EVENTS

$$\frac{Data}{MC_{(IU)}} = 0.55 \pm 0.04_{stat} \pm 0.06_{sys} \pm 0.14_{th}$$

$$\frac{Data}{MC_{(ID+UGS)}} = 0.70 \pm 0.04_{stat} \pm 0.07_{sys} \pm 0.18_{th}$$



— MC expectation (no oscillations)

- - - maximal mixing, $\Delta m^2 = 2.5 \times 10^{-3} \text{ eV}^2$

- Data consistent with a **constant deficit** in all **zenith angle bins** (IU: $\chi^2/\text{d.o.f.} = 3.1/4$ on shape)

UP-DOWN ASYMMETRY

$$\text{Ratio } R = \frac{\text{In Up}}{\text{In Down} + \text{UpG Stop}}$$

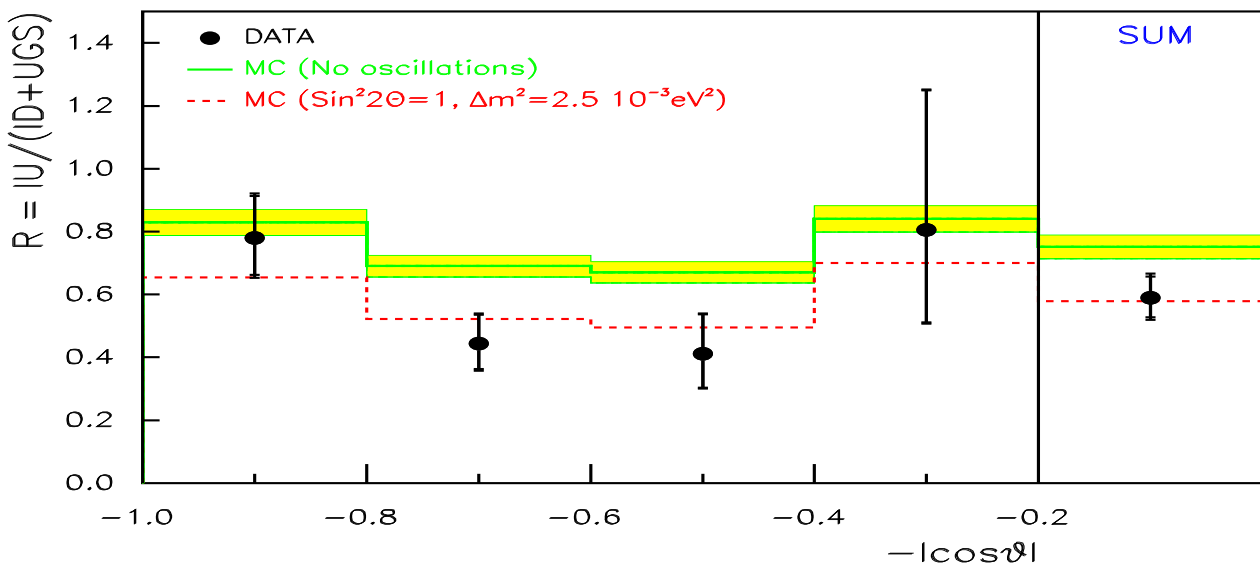
- Most of the theor. uncertainties canceled (<5%)
- Systematic errors reduced (~6%)

Data: $R = 0.59 \pm 0.07_{\text{stat}}$

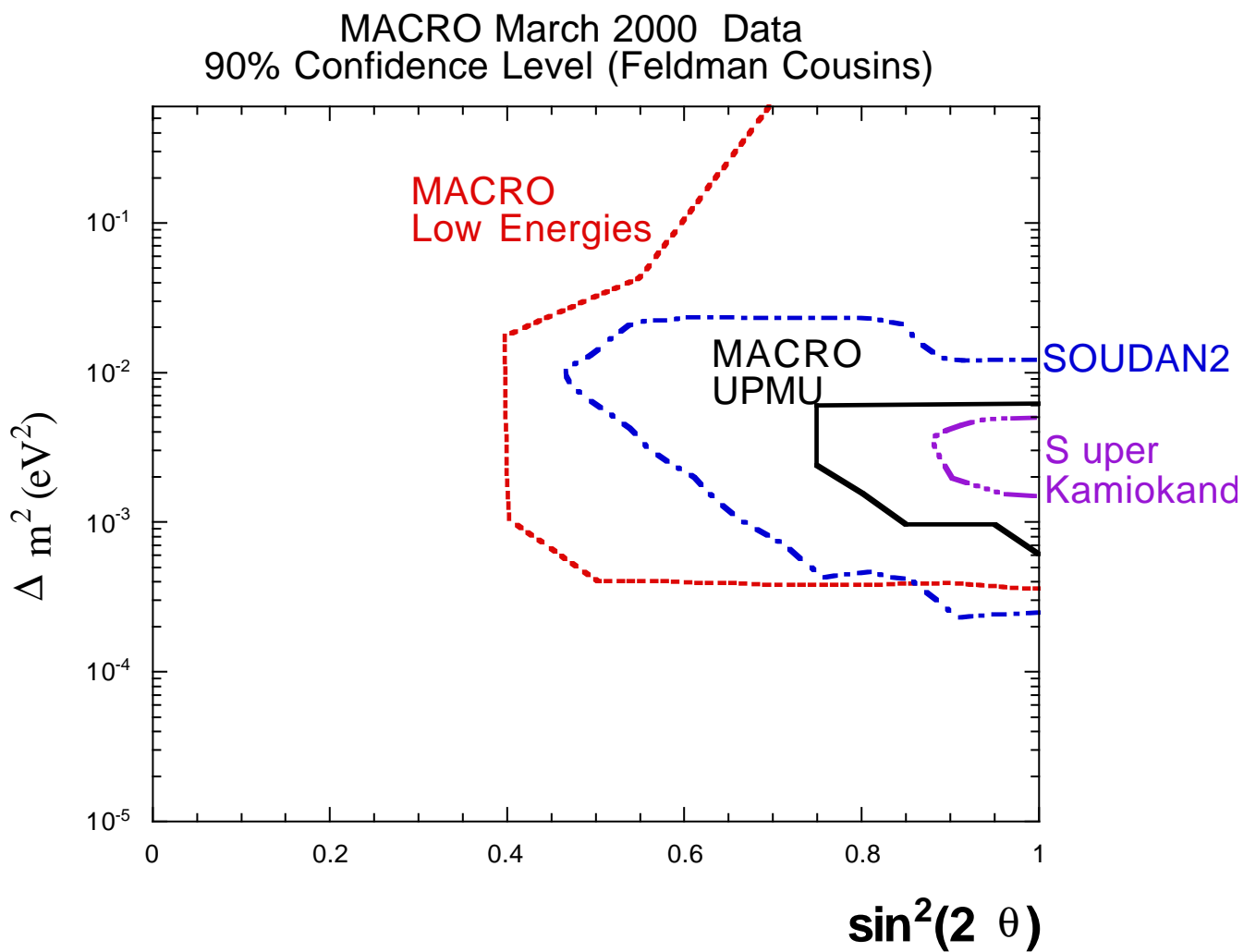
Expected (No oscillations): $R = 0.75 \pm 0.04_{\text{sys}} \pm 0.04_{\text{th}}$
 → compatibility ~ 2.7%

Expected with

$\nu_{\mu} \rightarrow \nu_{\tau}$ oscillations $R = 0.58 \pm 0.03_{\text{sys}} \pm 0.03_{\text{th}}$
 (maximal mixing and $\Delta m^2 = 2.5 \times 10^{-3} \text{ eV}^2$)



Confidence level regions ($\nu_{\mu} \rightarrow \nu_{\tau}$ oscillations)



Conclusions

1) High energy events

(Upward Through-Going Muons)

Angular distributions **more regular** than in the past
(Only statistics, no change in the data analysis in the last 3 years!)

$\chi^2 = 11.2/9$ d.o.f. for $\nu_\mu \rightarrow \nu_\tau$

Good agreement with the $\nu_\mu \rightarrow \nu_\tau$ oscillations with maximum mixing and Δm^2 around 0.0025 eV^2 :

2) Sterile Neutrino and Matter effects:

Two flavor sterile neutrino oscillations disfavored at > 98% respect $\nu_\mu \rightarrow \nu_\tau$ with maximum mixing (from the ratio test and 5% systematic error in each bin)

3) Low energy events:

event deficit, no zenith angle distortion,

Up-Down asymmetry ($\approx 3\%$) \implies

agreement with the $\nu_\mu \rightarrow \nu_\tau$ oscillations with maximum mixing and Δm^2 around 0.0025 eV^2 :